

## HINDU TRIGONOMETRY

by

BIBHUTIBHUSAN DATTA AND AVADHESH NARAYAN SINGH

Revised by

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### 1. TRIGONOMETRICAL FUNCTIONS. DEFINITIONS.

The Hindu name for the science of Trigonometry is *Jyotpatti-gaṇita* or “The science of calculation for the construction of the sine.”<sup>1</sup> It is found as early as in the *Brāhma-sphuṭa-siddhānta* of Brahmagupta (628).<sup>2</sup> Sometimes that name is simplified into *Jyā-gaṇita* (or “The science of calculation of the sines”).<sup>3</sup> In very recent years there has appeared the name *Trikonaṃiti*<sup>4</sup>, which is a literal as well as phonetic rendering of the Greek name for the science.

The Hindus introduced and usually employed three trigonometrical functions, namely *vyā*, *koṭi-jyā* and *utkrama-jyā*. It should be noted that they are functions of an arc of a circle, but not of an angle. If  $AP$  be an arc of a circle with centre at  $O$ , then

<sup>1</sup>*Jyā* (“sine”) + *utpatti* (“construction”, “generating”) + *gaṇita* (“the science of calculation”).

<sup>2</sup>*BrSpSI*, xii, 66.

<sup>3</sup>Compare *SITVI*, ii, 1.

<sup>4</sup>*Trikona* (“triangle”) + *ṃiti* (“measure”).

its  $jyā=PM$ ,  $koṭi-jyā=OM$  and  $utkrama-jyā=OA-OM=AM$ . Hence their relation with modern trigonometrical functions will be

$jyā AP=R\sin \theta$ ,  $koṭi-jyā AP=R\cos \theta$ ,  $utkrama-jyā AP=R-R\cos \theta=R\text{versin } \theta$ , where  $R$  is the radius of the circle and  $\theta$  the angle subtended at the centre by the arc  $AP$ . Thus the values of the Hindu trigonometrical functions vary with the radius chosen. The earliest Hindu treatise in which the above trigonometrical functions are now found recorded is the *Sūrya-siddhānta*.

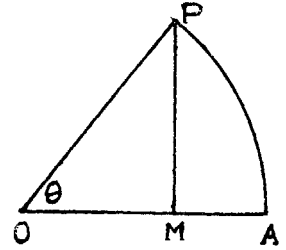


Fig. 1

*Jyā*

The Sanskrit word *jyā* means “a bow-string”; and hence “the chord of an arc”, for the arc is called “a bow” (*dhanu*, *cāpa*). Its synonyms are *jīvā*, *siñjini*,<sup>1</sup> *guṇa*, *maurvī*, etc. This trigonometrical function is also called *ardha-jyā*<sup>2</sup> (“half-chord”) or *jyārdha*<sup>3</sup> (“chord-half”). Thus Bhāskara II (1150) explicitly observes, “It should be known that *ardha-jyā* is here called “*jyā*”.”<sup>4</sup> Parameśvara (1430) remarks:

“A part of a circle is of the form of a bow, so it is called the “bow” (*dhanu*). The straight line joining its two extremities is the “bow-string” (*jīvā*); it is really the “full-chord” (*samasta-jyā*). Half of it is here (called) the “half-chord” (*ardha-jyā*), and half that arc is called the “bow” of that half-chord. In fact the *R*sine (*jyā*) and *R*cosine (*koṭi-jyā*) of that bow are always half-chords.”<sup>5</sup>

Kamalākara (1658) is more explicit. “Having seen the brevity”, says he, “the half-chords are called *Jyā* by mathematicians in this (branch of) mathematics and are used accordingly.”<sup>6</sup> The function *jyā* is sometimes distinguished as *krama-jyā*<sup>7</sup> or *kramārdha-jyā*,<sup>8</sup> from *krama*, “regular” or “direct” meaning “direct sine” or “direct half-chord”.

It may be noted that the modern term *sine* is derived from the Hindu name. The Sanskrit term *jīvā* was adopted by the early Arab mathematicians but was pronounced as *jība*. It was subsequently corrupted in their tongue into *jaib*. The latter word was confused by the early Latin translators of the Arabic works such as Gherardo of Cremona (c. 1150 A.D.) with a pure Arabic word of alike phonetism but meaning differently “bosom” or “bay” and was rendered as *sinus*, which also signifies “bosom” or “bay”.<sup>9</sup>

<sup>1</sup>*ŚiDVṛ*, ii. 9; *MSI*, iii. 2.

<sup>2</sup>*Ā*, i. 10; *BrSpSi*, ii. 2.

<sup>3</sup>*SūSi*, ii. 15; *Ā*, ii. 11, 12; *BrSpSi*, xxi. 17, 22.

<sup>4</sup>*ŚiSi*, *Graha*, iii. 2.

<sup>5</sup>*Ā*, ii. 11 (*Com.*).

<sup>6</sup>*ŚiTV*, ii. 52.

<sup>7</sup>*PSI*, iv. 28 (*Kramaśo jyā*); *BrSpSi*, ii. 15; vii. 12.

<sup>8</sup>*ŚiDVṛ*, ii. 1.

<sup>9</sup>*CF. Nouv. Ann. Math.*, XIII (1854), p. 393; Smith, *History II*, p. 616.

The degeneration and variations of the term *kramajyā* are still more interesting. In the Arabic tongue it was corrupted into *karaja* or *kardaja*. According to *Fihrist*, the title of a work of Ya'kūb ibn Ṭārik (c. 770 A.D.) is "On the table of *kardaja*." This table was copied from the *Brāhma-sphuṭa-siddhānta* of Brahmagupta. In the same connexion, al-Khowārizmī (825) used the variant *karaja*. In the Latin translations of the term we find several variants such as *kardaga*, *karkaya*, *gardaga* or *cardaga*. These terms had in foreign lands also the restricted uses for the arc of  $3^{\circ} 45'$ , sometimes of  $15^{\circ}$ .<sup>1</sup>

### *Koṭi-Jyā*

The Sanskrit word *koṭi* means, amongst others "the curved end of a bow" or "the end or extremity in general"; hence in Trigonometry it came to denote "the complement of an arc to  $90^{\circ}$ ."<sup>2</sup> So the radical significance of the term *koṭi-jyā* is "the *jyā* of the complementary arc". But it began early to be used as an independent technical term.<sup>3</sup> The modern term cosine appears to be connected with *koṭijyā*, for in Hindu works, particularly in the commentaries *koṭijyā* is often abbreviated into *kojyā*. When *jyā* became *sinus*, *kojyā* naturally became *ko-sinus* or *co-sinus*.

### *Utkrama-Jyā*

*Utkrama* means "reversed", "going out" or "exceeding". Hence the term *utkrama-jyā* literally means "reversed sine". This function is so called in contradistinction to *krama-jyā*, for it is, rather its tabular values are, derived from the tabular values of the latter by subtracting the elements from the radius in the reversed order. Or in other words it is the exceeding portion of the *krama-jyā* taken into consideration in the reversed order. Thus it is stated :

"The (tabular) versed sines are obtained by subtracting from the radius the (tabular) sines in the reversed order."<sup>4</sup>

"They (*jyārḍha*), (being subtracted from the radius), in the reversed order beginning from the end, will certainly give the versed sines, that is, the arrows."<sup>5</sup>

Again, it is noteworthy that from a table of differences of sines, the successive sines are obtained by adding the differences in the direct order (from the top) whereas the corresponding versed sines will be found by adding the elements in the reversed order (from the end). This fact has been particularly noted by Sūryadeva Yajvā (born 1191 A.D.) and Śrīpati (1039 A.D.). The former observes:

<sup>1</sup>Woepeke, F. "Sur le mot *kardaga* et sur une méthode indienne pour calcul les sinus", *Nouv. Ann. Math.*, XIII (1854), pp. 386-393; Braunmühl, A. *Geschichte der Trigonometrie*, 2 vols., Leipzig, 1900, 1903 (hereafter referred to as Braunmühl, *Geschichte*); Vol. I, pp. 44, 45, 78, 102, 110, 120; vide also Sarton's note on the point in *Isis*, xiv (1930), pp. 421f.

<sup>2</sup>In Hindu mathematics, the term *koṭi* also denotes "the side of a right-angled triangle."

<sup>3</sup>Compare *ŚIDVṛ*, ii. 30 (infra p. 10).

<sup>4</sup>*SūSi*, ii. 22.

<sup>5</sup>*BrSpSi*, xxi. 18; Compare also *MSi*, iii. 3.

“In order to get the direct sines (*krāma-jyā*), these (tabular) differences of sines (*khaṇḍa-jyā*) should be added regularly from the beginning; and in order to determine the reversed sines (*utkrāma-jyā*), they should be added in the reversed order from the end.”<sup>1</sup>

Śripati says:

“The difference of sines are called *vyākhaṇḍa* (tabular “difference of sines”); (adding them) in the reversed way beginning from the end will be obtained the versed sines (*vyasta-jyā*) of the half-arcs equal to the 96th parts of the celestial circle.”<sup>2</sup>

This function is also called *vyasta-jyā*<sup>3</sup> (from *vyasta*, “cast or thrown asunder”, “reversed”) or *viloma-jyā*<sup>4</sup> (from *viloma*, “reverse”). Occasionally it is termed *utkrāma-jyārḍha*.<sup>5</sup> Another name for it is “arrow” (*iṣu*, *bāṇa*).<sup>6</sup> Bhāskara II observes:

“What is really the arrow between the bow and the bowstring is known amongst the scholars here (i.e. in Trigonometry) as the versed sine.”<sup>7</sup>

So also says Kamalākara (1658):

“What lies between the chord and the arc, like the arrow, is the versed sine.”<sup>8</sup>

### *Tangent and Secant*

The Hindus approached very near the tangent and secant functions and actually employed them in astronomical calculations, though they did not expressly recognise them as separate functions. The *Sūrya-siddhānta* gives the following rule for calculating the equinoctial midday shadow of the gnomon at a station:

“The sine of the latitude (of the station) multiplied by 12 and divided by the cosine of the latitude gives the equinoctial mid-day shadow.”<sup>9</sup>

Here 12 is the usual height of a Hindu gnomon. So that

$$S = \frac{jyā \phi \times h}{kojyā \phi}$$

where  $\phi$  denotes the latitude of the place,  $S$ =equinoctial mid-day shadow and  $h$ =gnomon. This is equivalent to

$$S = h \tan \theta.$$

<sup>1</sup>*A* i. 10 (Com.).

<sup>2</sup>*SiSe*, xvi. 10.

<sup>3</sup>*BrSpSi*, ii. 5; *MSi*, iii. 3, 6.

<sup>4</sup>*SiDV*, I, ii. 5.

<sup>5</sup>*SūSi*, ii. 22, 27.

<sup>6</sup>*BrSpSi*, xxi. 18.

<sup>7</sup>*SiSi*, *Gola*, xiv. 5; Compare also *Graha*, ii. 20 (Gloss).

<sup>8</sup>*SiTV*, ii. 58.

<sup>9</sup>*SūSi*, iii. 16.

Again to find the mid-day shadow ( $s$ ) of the gnomon ( $h$ ) and the hypotenuse ( $d$ ), having known the meridian zenith distance ( $z$ ) of the sun, we have the rules:<sup>1</sup>

$$s = h \tan z, \quad d = h \sec z.$$

Similar rules occur in other astronomical works also.<sup>2</sup> In the *Gaṇita-sāra-saṃgraha* of Mahāvira (850) by the term “shadow” of a gnomon is sometimes meant the ratio of the actual shadow to the height of the gnomon.<sup>3</sup> This ratio, as has been just stated, is equal to the tangent of the zenith distance of the sun.

### Quadrants

A circle is ordinarily divided into four equal parts, called *vr̥tta-pāda*, by two perpendicular lines, usually the east-to-west line and the north-to-south line. The quadrants are again classified into odd (*ayugma*, *viṣama*) and even (*yugma*, *sama*). Earlier Hindu writers do not explain this fact fully and particularly. Thus Bhāskara I (629) simply observes: “Three signs form a quadrant”.<sup>4</sup> Lalla writes:

“Three anomalistic signs form a quadrant. The quadrants are successively distinguished as odd and even.”<sup>5</sup>

But the description of Bhāskara II (1150) is very full. He says:

“Three signs together form a quadrant. In a circle there will be four such; and they should be successively called odd and even.”<sup>6</sup>

He then explains it further thus:

“On a plane surface describe a circle of any specified radius with a pair of compasses. Mark on its circumference 360 degrees. Draw the east-to-west and north-to-south lines through its centre. These lines will divide the circle into quadrants, which should be taken into consideration in the leftwise manner (*savya-krama*, that is ‘anti-clockwise’) proceeding from the east-point (*prācī*); they should be called odd and even (quadrants) successively.”<sup>8</sup>

### Variation in value

As regards the variation in the value of a trigonometrical function as its argument changes, Bhāskara II observes as follows:

<sup>1</sup>*SūSi*, iii, 21.

<sup>2</sup>*PSi*, iv, 22.

<sup>3</sup>*GSS*, ix, 8½.

<sup>4</sup>*MBh*, iv, 1; *LBh*, ii, 1.

<sup>5</sup>*SiDVr*, ii, 10.

<sup>6</sup>*SiSi*, *Graha*, ii, 19.

<sup>7</sup>The Sanskrit term *savya-krama* ordinarily signifies the “clockwise direction”; but it may also denote the “anti-clockwise direction.”

<sup>8</sup>*SiSi*, *Graha*, ii, 19 (*Gloss*).

“In the first quadrant, mark a point on the circumference of the circle at any optional distance from the east point. The perpendicular distance of that point from the east-to-west line is called the *Rsine* (*doḥ-jyā*); and its distance from the north-to-south line is the *Rcosine* (*koṭi-jyā*). The corresponding arcs are called *bhuja* and *koṭi*. (Starting from the east point) as the point gradually moves forward in the same way (i.e. anti-clockwise), the *Rsine* increases and the *Rcosine* decreases. When the point arrives at the end of the quadrant, the *Rcosine* vanishes and the *Rsine* is equal to the radius. Then in the second quadrant, the *Rcosine* increases; at the end of that quadrant the *Rcosine* is maximum (irrespective of sign) and the *Rsine* vanishes.”<sup>1</sup>

One fact perhaps deserves a particular notice here. It is that in Hindu trigonometry the *jyā* of an arc of  $90^\circ$  in a circle is equal to the radius of that circle. On account of that, the radius is called in Hindu mathematics by the terms *tri-jyā*, *tri-bha-jyā*, *tribhavana-jyā*, etc., every one of which literally means the “sine of three signs.” The radius is also called *viṣkambhārḍha*, *vyāsārḍha*, or *ardha-vyāyāma* meaning the “semi-diameter”. All these terms are very old.<sup>2</sup>

#### *Functions of a complement or supplement*

*Sūrya-siddhānta* says:

“In odd quadrants, the arc passed over gives the *Rsine*, while the arc to be passed over gives the *Rcosine*; and in the even quadrants, the arc to be passed over gives the *Rsine* and that passed over gives the *Rcosine*.”<sup>3</sup>

Bhāskara I writes:

“In the odd quadrants the arc described and that to be described should respectively be known as the *bhuja* and *koṭi*; but in the even quadrants they are respectively the *koṭi* and *bhuja*; this is the fact.”<sup>4</sup>

Lalla remarks:

“When (the anomaly<sup>5</sup> is) greater than  $90^\circ$ , it is subtracted from the semi-circle (i.e.  $180^\circ$ ); when greater than the semi-circle,  $180^\circ$  is subtracted from it; when greater than  $270^\circ$ , it is subtracted from the complete circle (i.e.  $360^\circ$ ); the remainder is called the (corresponding) *bhuja* by the expert in the subject.”<sup>6</sup>

<sup>1</sup>*Ibid.*, ii. 20 (*Gloss*). The Sanskrit terms *jyā* and *koṭyā* have been translated as *Rsine* and *Rcosine* because they are equal to  $R \times \text{sine}$  and  $R \times \text{cosine}$  respectively.

<sup>2</sup>Compare *ApŚiSū*, vii. 11 (*arāha-vyāyāma*); *Jambūdvīpasamāsa* of *Umāsvāti*, iv (*vyasārḍha*); *Tattvārthādhiḡama-sūtra-bhāṣya*, iv. 14 (*viṣkambhārḍha*).

<sup>3</sup>*SūSi*, ii. 30.

<sup>4</sup>*LBh*, ii. 1-2; Compare also *MBh*, iv. 8-9.

<sup>5</sup>It is in connection with the treatment of the anomaly that the remark of Lalla, as of several other Hindu mathematicians, occurs.

<sup>6</sup>*ŚiDVr*, ii. 10-11.

In the words of Brahmagupta:

“The *Rsine* and *Rcosine* (are obtained) in the odd quadrants from the arc passed over and to be passed over (respectively); and in the even quadrants in the reverse way.”<sup>1</sup> Or,

“In the odd quadrants (the *Rsine* is determined) from the arc described and in the even quadrants from the arc to be described.”<sup>2</sup>

“(For the determination of) the *Rsine* (proceed with the anomaly as it is) when the anomaly is less than three signs (i.e.  $90^\circ$ ); when greater than three signs subtract it from six signs; when greater than six signs, subtract six signs (from it); when greater than nine signs, subtract it from the complete circle.”<sup>3</sup>

Mañjula (932) says:

“In the odd quadrants, the *bhuja* and *koṭi* are (to be calculated) from the arc described and that to be described (respectively); but in the even quadrants in the contrary way.”<sup>4</sup>

His commentator and younger contemporary Praśastidhara (962) dilates upon this point thus:

“In the odd quadrant, where the anomaly is less than three signs (i.e.  $90^\circ$ ), the *Rsine* should be calculated from it and the *Rcosine* should be calculated after subtracting that from  $90^\circ$ . In the even quadrant, where the anomaly exceeds  $90^\circ$  but is less than  $180^\circ$ ; in that case the *Rsine* should be taken after subtracting it from  $180^\circ$  and the cosine after subtracting  $90^\circ$  from it. In the odd quadrant, where the anomaly is greater than  $180^\circ$ , but less than  $270^\circ$ , the *Rsine* should be calculated after subtracting  $180^\circ$  from it and the *Rcosine* after subtracting it from  $270^\circ$ . In the even quadrant when the anomaly exceeds  $270^\circ$ , but is less than  $360^\circ$ , the *Rsine* is determined after subtracting it from  $360^\circ$ , and the *Rcosine* after subtracting  $270^\circ$  from it.”<sup>5</sup>

Śripati (c. 1039) remarks:

“In the odd and even quadrants, the arc passed over and to be passed over (respectively) is the *bhuja* and the *koṭi* is otherwise. Or, as the learned have said, the *Rsine* of  $90^\circ$  minus the anomaly is the *Rcosine* (of the anomaly).”<sup>6</sup>

<sup>1</sup>*BrSpSi*, ii. 12.

<sup>2</sup>*KK*, I, i. 16.

<sup>3</sup>*KK*, I, i. 16.

<sup>4</sup>*LMā*, ii. 2.

<sup>5</sup>Commentary on the same.

<sup>6</sup>*SiSe*, iii. 13.

And Bhāskara II:

“In the odd quadrants, the arc passed over and in the even quadrants the arc to be passed is the *bhuja*. Ninety degrees minus the *bhuja* is said to be the *koṭi*.”<sup>1</sup>

The above results can be represented graphically thus :

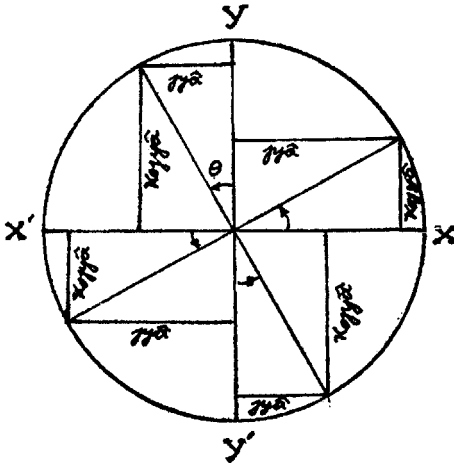


Fig. 2

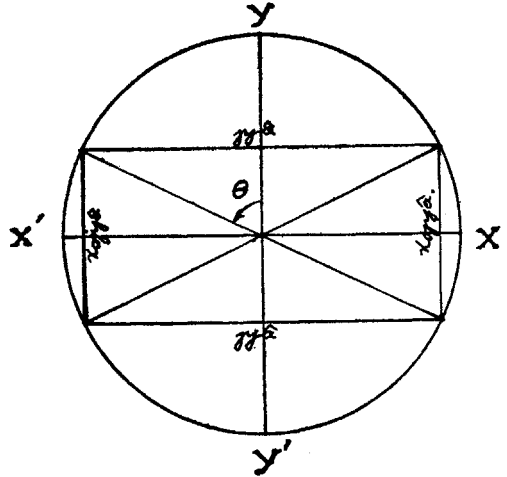


Fig. 3

Relation between Functions

Varāhamihira says:

“The Rsine of 90° minus latitude is the Rcosine of the latitude.”<sup>2</sup>

Lalla:

“The square of the base-sine (*bhuja-jyā*) is subtracted from the square of the radius; the square root of the remainder is the Rcosine; or it is the Rsine of 90° minus the *bhuja* arc.”<sup>3</sup>

$$\sqrt{R^2 - (jyā \alpha)^2} = kojyā \alpha$$

or  $kojyā \alpha = jyā (90^\circ - \alpha)$

where *kojyā* is the usual Hindu symbol for *koṭi-jyā*.

<sup>1</sup>SiŚi, *Graha*, ii. 19.

<sup>2</sup>PSi. iv. 28.

<sup>3</sup>SiDV<sub>r</sub>, ii. 30.



Brahmagupta says:

“The radius diminished by the versed *Rsine* of an arc or of its complement will give the *Rsine* of the other. The square-root of the difference of the square of the radius and that of the *Rsine* of an arc or of its complement will be the *Rsine* of the other.”<sup>1</sup>

$$\begin{aligned} R - utjy\bar{a} \alpha &= jy\bar{a} (90^\circ - \alpha), \\ R - utjy\bar{a} (90^\circ - \alpha) &= jy\bar{a} \alpha, \\ \sqrt{R^2 - (jy\bar{a} \alpha)^2} &= jy\bar{a} (90^\circ - \alpha), \\ \sqrt{R^2 - \{jy\bar{a}(90^\circ - \alpha)\}^2} &= jy\bar{a} \alpha \end{aligned}$$

where *utjyā* is the usual abbreviation for *utkrama-jyā*.

“The direct *Rsine* of the excess of an arc over  $90^\circ$  added to the radius will give versed *Rsine* of that arc.”<sup>2</sup>

$$R + jy\bar{a} (\alpha - 90^\circ) = utjy\bar{a} \alpha,$$

where  $\alpha > 90^\circ$ .

Śrīpati writes:

“The square of the radius is diminished by the square of the *Rsine*; the square-root of the remainder will be the *Rcosine*. Again the square-root of the square of the radius minus the square of the *Rcosine* will be the *Rsine*. The radius minus the versed *Rsine* of the complement of an arc is equal to the *Rsine* of the arc, and minus the versed *Rsine* of the arc becomes the *Rsine* of the other (i.e. complement).”<sup>3</sup>

The treatment of Bhāskara II is exhaustive. He says:

“Subtract from the radius the direct *Rsine* of an arc and of its complement; the results will be the versed *R sines* of the complement and the arc (respectively). Subtract from the radius the versed *Rsine* of an arc and of its complement; the remainders will be the direct *Rsines* of the complement and the arc (respectively).”<sup>4</sup>

$$\begin{aligned} R - jy\bar{a} \alpha &= utjy\bar{a} (90^\circ - \alpha), \\ R - jy\bar{a} (90^\circ - \alpha) &= utjy\bar{a} \alpha, \\ R - utjy\bar{a} \alpha &= jy\bar{a} (90^\circ - \alpha), \\ R - utjy\bar{a} (90^\circ - \alpha) &= jy\bar{a} \alpha. \end{aligned}$$

<sup>1</sup>*BrSpSi*, xiv. 7.

<sup>2</sup>*Ibid*, vii. 12.

<sup>3</sup>*SiSe*, iii. 14.

<sup>4</sup>*SiŚī*, *Graha*, ii. 20; also *Gola*, v. 2; xiv. 5.

“The square of the *Rsine* of an arc and of its complement are (severally) subtracted from the square of the radius, the square-roots of the results are (respectively) the *Rsines* of the complement and of the arc.”<sup>1</sup>

$$\sqrt{R^2 - (jy\bar{a} \alpha)^2} = jy\bar{a} (90^\circ - \alpha); \quad \sqrt{R^2 - \{jy\bar{a} (90^\circ - \alpha)\}^2} = jy\bar{a} \alpha.$$

“The square of the radius is diminished by the square of the *Rsine* of an arc; the square-root of the result is the *Rcosine* of the arc.”<sup>2</sup>

$$\sqrt{R^2 - (jy\bar{a} \alpha)^2} = kojy\bar{a} \alpha.$$

Kamalākara writes:

“The square-root of the square of the radius diminished by the square of the *Rsine* of an arc, is the *Rcosine* of the arc; similarly, the square-root of the square of the radius diminished by the *Rcosine* of an arc, is the *Rsine* of the arc. Again, the *Rsines* of an arc and its complement when subtracted from the radius will give the versed *Rsines* of the complement and the arc (respectively).”<sup>3</sup>

$$\sqrt{R^2 - (jy\bar{a} \alpha)^2} = kojy\bar{a} \alpha, \quad \sqrt{R^2 - (kojy\bar{a} \alpha)^2} = jy\bar{a} \alpha, \\ R - jy\bar{a} \alpha = utjy\bar{a} (90^\circ - \alpha), \quad R - kojy\bar{a} \alpha = utjy\bar{a} \alpha.$$

#### *Change of Sign of a Function*

The Hindus were fully aware of the changes of sign of a trigonometrical function according as its argument lies in different quadrants. Though nowhere do we find any systematic treatment of this principle in any Hindu work there are ample concrete instances of its application in almost all their important astronomical treatises. Thus it is stated in the *Sūrya-siddhānta*:

“The *śighra koṭiphala* is positive, when the *kendra* (mean anomaly) lies in a position beginning with the Capricorn; and it is to be subtracted from the radius in a position beginning with the Cancer.”<sup>4</sup>

Now according to the *Sūrya-siddhānta* and other Hindu astronomical works, the *śighra koṭiphala* (the result derived from the complement of the distance from the conjunction) is given by  $D \cos \theta$ , where  $\theta$  is the *śighra kendra*<sup>5</sup> (the distance of the mean planet from its apex of swiftest motion; hence mean *śighra* anomaly) and  $D$ , a certain known constant. The Cancer is the fourth sign of the Zodiac and Capricorn is the tenth sign. Again the motion of the mean planet is anti-clockwise. Hence it is clear from

<sup>1</sup> *SiŚi, Graha*, ii. 21.

<sup>2</sup> *SiŚi, Gola*, v. 2; xiv. 4.

<sup>3</sup> *SiTVi*, ii. 56-7.

<sup>4</sup> *SūSi*, ii. 40.

<sup>5</sup> “Subtract the longitude of a planet from that of its apex of slowest motion (*mandocca*); so also subtract it from that of its apex of swiftest motion (conjunction); the result (in either case) is its *Kendra*.” *SūSi*, ii. 29.

the above rule that the author was aware that the cosine of an angle lying between  $0^\circ$  and  $90^\circ$  or between  $270^\circ$  and  $360^\circ$  is positive and that it is negative when the angle lies between  $90^\circ$  and  $270^\circ$ .

Again it has been said:

“In case of the *manda* and *śighra* corrections of all planets, the *phala* (equation) will be positive, if the *kendra* lies in the six signs beginning with the Aries and it will be negative in the six signs beginning with the Libra.”<sup>1</sup>

Now the *phala* is defined as  $\text{arc}(D' \sin \theta)$ , where  $D'$  does not change sign. Hence clearly the author knows that the sign is positive in the first two quadrants and negative in the other two quadrants.

Similar rules are found in other treatises of astronomy.<sup>2</sup> The statement of Mañjula (932) is more explicit and fuller. He says:

“The (mean) planet when diminished by its apogee or aphelion is the *kendra* (mean anomaly). Its *Rsine* is positive or negative in the upper or lower halves (of the quadrants); and its *Rcosine* is positive, negative, negative, and positive (respectively) according to the (successive) quadrants.”<sup>3</sup>

Thus the Hindus knew very early what in modern trigonometrical notations will be expressed as—

$$\sin(\pi \mp \theta) = \pm \sin \theta, \quad \cos(\pi \mp \theta) = -\cos \theta$$

$$\sin(2\pi - \theta) = -\sin \theta, \quad \cos(2\pi - \theta) = +\cos \theta$$

$$\sin\left(\frac{\pi}{2} \mp \theta\right) = +\cos \theta, \quad \cos\left(\frac{\pi}{2} \mp \theta\right) = \pm \sin \theta$$

$$\sin\left(\frac{3\pi}{2} \mp \theta\right) = -\cos \theta, \quad \cos\left(\frac{3\pi}{2} \mp \theta\right) = \mp \sin \theta$$

Again it has been stated before that according to a rule of Brahmagupta

$$R + jyā(\alpha - 90^\circ) = utjyā \alpha.$$

But by definition,

$$utjyā \alpha = R - kojyā \alpha = R - jyā(90^\circ - \alpha).$$

<sup>1</sup>*SūSi*, ii. 45.

<sup>2</sup>For instance *Ā*, iii. 22; *MBh*, iv. 5, 9; *LBh*, ii. 6; *ŚiDVṛ*, ii. 32; *BrSpSi*, ii. 14ff.

<sup>3</sup>*LMā*, ii. 1.

These clearly show that the author knows that the value of the sine function changes sign along with its argument.

Or symbolically

$$\sin(\pm\theta) = \pm\sin\theta.$$

## 2. TRIGONOMETRICAL FORMULAE

$$(1) \sin^2\theta + \cos^2\theta = 1$$

It has been stated before that according to Hindu astronomers, if  $\alpha$  be an arc of a circle of radius  $R$

$$\sqrt{R^2 - (jy\bar{\alpha})^2} = kojy\bar{\alpha}, \quad \sqrt{R^2 - (kojy\bar{\alpha})^2} = jy\bar{\alpha}.$$

These are of course equivalent to the modern formulae

$$\sqrt{1 - \sin^2\theta} = \cos\theta, \quad \sqrt{1 - \cos^2\theta} = \sin\theta$$

$$\text{or } \sin^2\theta + \cos^2\theta = 1,$$

where  $\theta$  is the angle subtended at the centre of the circle by the arc  $\alpha$ .

$$(2) 4 \sin^2 \frac{\theta}{2} = \sin^2\theta + \text{versin}^2\theta$$

This formula has been stated first by Varāhamihira (505).

He says:

“To find the Rsine of any other desired arc, double the arc and subtract from the quarter of a circle; diminish the radius by the Rsine of the remainder. The square of half the result is added to the square of half the Rsine of the double arc. The square-root of the sum is the desired Rsine.”<sup>1</sup>

Let the arc  $XP = \text{arc } PQ = \alpha$ ; then arc  $QY = 90^\circ - 2\alpha$ .

Now

$$\begin{aligned} XQ^2 &= QT^2 + TX^2 \\ \text{or } 4XD^2 &= QT^2 + TX^2 \\ \text{or } XD^2 &= (QT/2)^2 + (TX/2)^2. \end{aligned}$$

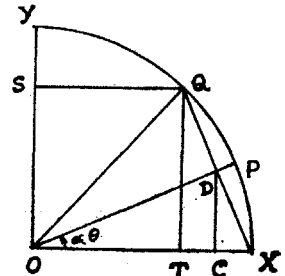


Fig. 4

<sup>1</sup>PSI, iv. 2f.

Hence

$$(jyā \alpha)^2 = \left( \frac{jyā 2\alpha}{2} \right)^2 + \left( \frac{R - jyā (90^\circ - 2\alpha)}{2} \right)^2;$$

which is equivalent to

$$4 \sin^2 \theta = \sin^2 2\theta + \text{versin}^2 2\theta.$$

Āryabhaṭa I (499) seems to have been aware of this formula before Varāhamihira. It reappears also in later works.

Brahmagupta says:

“The sum of the squares of the *Rsine* and versed *Rsine* of the same arc is divided by four; subtract this quotient from the square of the radius. Take the square-root of the two results. The former will be the *Rsine* of half that arc, and the other the *Rsine* of the arc equal to the quarter circle less that half.”<sup>1</sup>

The formula has been described almost similarly by Śrīpati (1039)<sup>2</sup>. Bhāskara II (1150) writes very briefly thus:

“Half the square-root of the sum of the square of the *Rsine* and of the versed *Rsine* of an arc, will be the *Rsine* of half that arc.”<sup>3</sup>

Parameśvara (1430) says:

“The square-root of the sum of the square of the *Rsine* and of the versed *Rsine* of an arc is the ‘whole chord’ (*samasta-jyā*) of that arc. Half that is the half-chord (i.e. the *Rsine*) of half that arc.”<sup>4</sup>

$$(3) \quad 2 \sin^2 \frac{\theta}{2} = 1 - \cos \theta$$

This is given first by Varāhamihira. He says:

“Twice any desired arc is subtracted from three signs (i.e. 90°), the *Rsine* of the remainder is subtracted from the *Rsine* of three signs. The result multiplied by sixty is the square of the *Rsine* of that arc.”<sup>5</sup>

<sup>1</sup> *Br. Sp. Si.*, xxi. 20f.

<sup>2</sup> *Si Śe.*, xvi. 14-5.

<sup>3</sup> *Si Śi.*, *Gola*, v. 4; xiv. 10.

<sup>4</sup> Quoted in his commentary of *Ā.*, ii. 11.

<sup>5</sup> *PSi.*, iv. 5.

In the fig. 4, page 50, since the triangles  $XCD$  and  $XDO$  are similar, we have:

$$XD : XC :: XO : XD$$

$$\therefore XD^2 = XO \cdot XC = \frac{1}{2} XO \cdot XT$$

$$\text{Hence } (jy\bar{a} \alpha)^2 = \frac{1}{2} R \{R - jy\bar{a} (90^\circ - 2\alpha)\}.$$

The factor  $\frac{1}{2} R$  on the right-hand side has been stated by Varāhamihira as 60 since he has taken the value of the radius to be equal to 120. In modern notations, the above formula becomes—

$$\sin^2 \theta = \frac{1}{2} (1 - \cos 2\theta).$$

This also follows easily from the preceding formula.

Brahmagupta says:

“The square-root of the fourth part of the versed  $R$ sine of an arc multiplied by the diameter is the  $R$ sine of half that arc.”<sup>1</sup>

Bhāskara II writes:

“Or, the square-root of half the product of the radius and the versed  $R$ sine of an arc, will be the  $R$ sine of half that arc.”<sup>2</sup>

He has further given the following proof of it.<sup>3</sup>

$$\text{Since } kojy\bar{a} \alpha = R - utjy\bar{a} \alpha$$

$$\text{so that squaring } (kojy\bar{a} \alpha)^2 = R^2 + (utjy\bar{a} \alpha)^2 - 2R \cdot utjy\bar{a} \alpha$$

$$\text{Therefore } R^2 - (kojy\bar{a} \alpha)^2 = 2R \cdot utjy\bar{a} \alpha - (utjy\bar{a} \alpha)^2$$

$$\text{Or, } (jy\bar{a} \alpha)^2 = 2R \cdot utjy\bar{a} \alpha - (utjy\bar{a} \alpha)^2$$

$$(jy\bar{a} \alpha)^2 + (utjy\bar{a} \alpha)^2 = 2R \cdot utjy\bar{a} \alpha$$

But by the formula (2), the righthand side is equal to

$$4 \left( jy\bar{a} \frac{\alpha}{2} \right)^2.$$

$$\text{Hence, } jy\bar{a} \frac{\alpha}{2} = \sqrt{\frac{1}{2} R \cdot utjy\bar{a}}.$$

This rule of Bhāskara II together with his proof has been reproduced by Kamalākara.<sup>4</sup>

<sup>1</sup>*BrSpSi*, xxi. 23.

<sup>2</sup>*SiŚi, Gola*, v. 5; xiv. 10.

<sup>3</sup>*SiŚi, Gola* (Gloss).

<sup>4</sup>*SiTVi*, ii. 78 and its commentary.

$$(4) \sin \frac{1}{2} (90^\circ \pm \theta) = \sqrt{\frac{1}{2} (1 \pm \sin \theta)}$$

This formula first appears in the works of Āryabhaṭa II (950). He says:

“The *Rsine* of any arc multiplied by the radius is subtracted from or added to the square of the maximum value of the *Rsine*; the square-root of half the results are extracted. These will be the *Rsine* of 45° decreased or increased by half that arc.”<sup>1</sup>

Let the arc  $XP$  be denoted by  $\alpha$ . Bisect the complementary arc  $YP$  at  $Q$ . Then

$$\begin{aligned} YP^2 &= YN^2 + NP^2 \\ &= (OY - PM)^2 + PN^2 \\ &= OY^2 + PM^2 + OM^2 - 2 OY \cdot PM. \end{aligned}$$

Therefore,

$$4 PC^2 = 2 (OP^2 - OP \cdot PM).$$

$$\text{Hence, } jy\bar{a} \frac{1}{2} (90^\circ - \alpha) = \sqrt{\frac{1}{2} (R^2 - R \cdot jy\bar{a} \alpha)}.$$

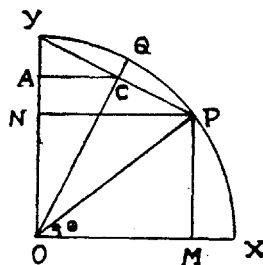


Fig. 5

Similarly it can be proved that

$$jy\bar{a} \frac{1}{2} (90^\circ + \alpha) = \sqrt{\frac{1}{2} (R^2 + R \cdot jy\bar{a} \alpha)}.$$

These are of course equivalent to

$$\sin \frac{1}{2} (90^\circ \pm \theta) = \sqrt{\frac{1}{2} (1 \pm \sin \theta)}.$$

Bhāskara II (1150) writes:

“The square of the radius is diminished or increased by the product of the radius and the *Rsine* of an arc; the square-root of half the results will be the *Rsine* of the half of 90° minus or plus that arc.”<sup>2</sup>

Kamalākara defines:

“The product of the radius and the *Rsine* of an arc is added to or subtracted from the square of the radius. The square-root of the half of the results are taken. They will respectively be the *Rsine* of the half of three signs plus or minus the arc.”<sup>3</sup>

<sup>1</sup>*MSi*, iii. 2.

<sup>2</sup>*SiSi, Gola*, xiv. 12.

<sup>3</sup>*SiTVi*, ii. 93

He adduces the following proof of it:<sup>1</sup>

$$R \pm jy\bar{a} \alpha = utjy\bar{a} (90^\circ \pm \alpha)$$

Squaring and adding  $\{jy\bar{a} (90^\circ \pm \alpha)\}^2$  to both the sides, we get  
 $R^2 + (jy\bar{a} \alpha)^2 + \{jy\bar{a} (90^\circ \pm \alpha)\}^2 \pm 2R jy\bar{a} \alpha = \{jy\bar{a} (90^\circ \pm \alpha)\}^2 + \{utjy\bar{a} (90^\circ \pm \alpha)\}^2$   
 or,  $2(R^2 \pm R jy\bar{a} \alpha) = 4 \{jy\bar{a} \frac{1}{2} (90^\circ \pm \alpha)\}^2$ , by formulae (1) and (2).

$$(5) \quad 2 \cos^2 \frac{\theta}{2} = 1 + \cos \theta$$

Bhāskara II remarks that if the arc  $\alpha$  in the formula

$$jy\bar{a} \frac{1}{2} (90^\circ \pm \alpha) = \sqrt{\frac{1}{2}(R^2 \pm R jy\bar{a} (90^\circ - \alpha))}$$

be substituted by its complement  $90^\circ - \alpha$ , it will still be true.<sup>2</sup>

So that,  $jy\bar{a} \frac{1}{2} (90^\circ \pm 90^\circ - \alpha) = \sqrt{\frac{1}{2}\{R^2 \pm R jya(90^\circ - \alpha)\}}$

which leads to,  $2 \cos^2 \frac{\theta}{2} = 1 + \cos \theta$ ,  $2 \sin^2 \frac{\theta}{2} = 1 - \cos \theta$ .

Kamalākara says:

“Half the *R*cosine of an arc is added to the *R*sine of one sign (i.e.  $30^\circ$ ) and the sum is multiplied by the radius; the square-root of the product should be known by the intelligent as the *R*cosine of half that arc.”<sup>3</sup>

$$kojy\bar{a} \frac{\alpha}{2} = \sqrt{R (jy\bar{a} 30^\circ + \frac{1}{2} kojy\bar{a} \alpha)}$$

or  $\cos^2 \frac{\theta}{2} = \sin 30^\circ + \frac{1}{2} \cos \theta = \frac{1}{2}(1 + \cos \theta)$

$$(6) \quad \sin^2 (45^\circ - \theta) = \frac{1}{2} (\cos \theta - \sin \theta)^2$$

Bhāskara II says:

“The square of the difference of the *R*sine and *R*cosine of an arc is halved; the square-root of the result is equal to the *R*sine of half the difference between that arc and its complement.”<sup>4</sup>

<sup>1</sup> *SiTVi*, (Gloss).

<sup>2</sup> *SiSi*, *Gola*, xiv. 12.

<sup>3</sup> *SiTVi*, ii. 91.

<sup>4</sup> *SiSi*, *Gola*, xiv. 14.



Denote the arc  $XP$  by  $\alpha$  ; cut off the arc  $YQ$  equal to the arc  $XP$ . Bisect the chord  $PQ$  by the point  $D$ .

Then,  $CP = PN - CN = PN - QS$   
 $= \text{kojyā } \alpha - jyā \alpha$   
 $= QT - PM = CQ.$

Therefore,  $PQ^2 = 2 CP^2$   
 $PD^2 = \frac{1}{2} CP^2$

or,  $jyā \frac{1}{2} (90^\circ - \alpha - \alpha) = \sqrt{\frac{1}{2} (\text{kojyā } \alpha - jyā \alpha)^2}$

which is equivalent to  
 $\sin (45^\circ - \theta) = \sqrt{\frac{1}{2} (\cos \theta - \sin \theta)^2}.$

Kamalākara writes :

“The *Rsine* of half the difference between an arc and its complement should be known by the intelligent in this (science) as equal to the square-root of half the square of the difference of the *Rsine* of the arc and of its complement.”<sup>1</sup>

His proof of the formula is substantially the same as that stated above.

(7)  $\cos 2 \theta = 1 - 2 \sin^2 \theta$

Bhāskara II gives:

“The square of the *Rsine* of an arc is divided by half the radius; the difference between this quotient and the radius is equal to the *Rsine* of the difference between that arc and its complement.”<sup>2</sup>

In the fig. 4 on page 50

$$QX^2 = QT^2 + TX^2 = QT^2 + (OX - OT)^2$$

$$= QT^2 + OT^2 + OX^2 - 2 OX \cdot OT.$$

or  $4 XD^2 = 2 OX^2 - 2 OX \cdot QS.$

Hence,  $QS = OX - \frac{XD^2}{OX/2}$

So that  $jyā(90^\circ - 2\alpha) = R - \frac{(jyā\alpha)^2}{R/2}$

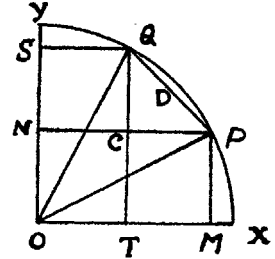


Fig. 6

<sup>1</sup>*Sitavi*, ii. 95.  
<sup>2</sup>*SiSi, Gola*, xiv. 15.

which is the same as

$$\cos 2\theta = 1 - 2 \sin^2 \theta.$$

This formula is practically the same as (3). In the words of Kamalākara :

“Twice the square of the *Rsine* of an arc is divided by the radius, the quotient is subtracted from the radius; the remainder will be the *Rsine* of the difference of the arc and its complement.”<sup>1</sup>

$$(8) \sin^2 \theta + \text{versin}^2 \theta = 2 \text{versin } \theta$$

$$(9) 2 \sin \theta \cos \theta + [\text{versin } \theta - \text{versin } (90^\circ - \theta)]^2 = 1$$

$$(10) (1 + \sin \theta) \cdot \text{versin } (90^\circ - \theta) = \cos^2 \theta$$

$$(11) 2 \sin \theta \pm [\text{versin } \theta \sim \text{versin } (90^\circ - \theta)] \\ = \sqrt{2 - [\text{versin } \theta \sim \text{versin } (90^\circ - \theta)]^2},$$

according as  $\sin \theta \lesseqgtr \cos \theta$

$$(12) (\cos \theta + \sin \theta)^2 + [\text{versin } \theta \sim \text{versin } (90^\circ - \theta)]^2 = 2.$$

Formulae (8) to (12) and similar others occur in the *Vaṭeśvara-siddhānta* of Vaṭeśvara (904).

### 3. ADDITION AND SUBTRACTION THEOREMS

Bhāskara II (1150) says:

“The *Rsin*es of any two arcs of a circle are reciprocally multiplied by their *Rcos*ines; the products are then divided by the radius; the sum of the quotients is equal to the *Rsine* of the sum of the two arcs; and their difference is the *Rsine* of the difference of the arcs.”<sup>2</sup>

If  $\alpha$  and  $\beta$  be any two arcs, then the rule says:

$$jyā (\alpha \pm \beta) = \frac{jyā \alpha \cdot kojyā \beta}{R} \pm \frac{kojyā \alpha \cdot jyā \beta}{R}$$

which is equivalent to

$$\sin (\theta \pm \phi) = \sin \theta \cos \phi \pm \cos \theta \sin \phi.$$

<sup>1</sup>*SITVi*, ii. 96.

<sup>2</sup>*SiŚi, Gola*, xiv. 21f.

In the words of Kamalākara (1658):

“The quotients of the *Rsines* of any two arcs of a circle divided by its radius are reciprocally multiplied by their *Rcosines*; the sum and difference of them (the products) are equal to the *Rsine* of the sum and difference respectively of the two arcs.”<sup>1</sup>

The rule for finding the *Rcosine* of the sum and difference of two arcs of a circle is enunciated by Kamalākara thus:

“The product of the *Rcosines* and of the *Rsines* of two arcs of a circle are divided by its radius; the difference and sum of them (the quotients) are equal to the *Rcosine* of the sum and difference (respectively) of the two arcs.”<sup>2</sup>

$$kojyā (\alpha \pm \beta) = \frac{kojyā \alpha \cdot kojyā \beta \mp jyā \alpha \cdot jyā \beta}{R}$$

which is equivalent to

$$\cos (\theta \pm \phi) = \cos \theta \cos \phi \pm \sin \theta \sin \phi.$$

Though we do not find this *Rcosine* theorem in the printed editions of the works of Bhāskara II, we are quite sure that it was known to him. For it has been attributed to him by his most relentless critic Kamalākara<sup>3</sup> as well as by his commentator Munīśvara.

The above theorems can be proved by methods algebraical as well as geometrical. Several such proofs were given by previous writers, observes Kamalākara<sup>4</sup> (1658). Unfortunately we have not been able to trace them as yet. The following two geometrical proofs are found in the *Siddhānta-tattva-viveka*<sup>5</sup> of Kamalākara.

*First Proof.* Let the arc  $YP = \beta$ , and arc  $YQ = \alpha$ ;  $\alpha$  being greater than  $\beta$ . Join  $OP, OQ$ .

Draw  $PN, PM$  perpendicular to  $OY, OX$  respectively. Also draw  $QS$  perpendicular to  $OY$ , and produce it to meet the circle again at  $Q'$ . Draw  $QT, Q'T'$  perpendicular to  $OP$ . Then  $PN = jyā \beta$ ,  $ON = kojyā \beta$ ,  $QS = jyā \alpha$ ,  $OS = kojyā \alpha$ ;  $PG = kojyā \beta - kojyā \alpha$ ,  $QG = jyā \alpha + jyā \beta$ ,  $QT = jyā (\alpha + \beta)$ ,  $PT = R - kojyā (\alpha + \beta)$ .

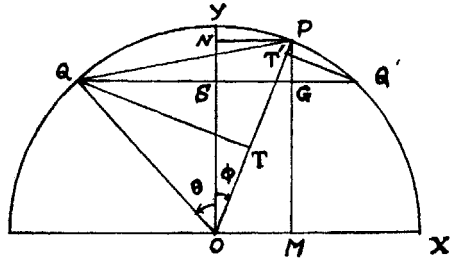


Fig. 7

<sup>1</sup>*SitVi*, ii. 68.

<sup>2</sup>*Ibid.*, ii. 69.

<sup>3</sup>Kamalākara remarks:

“*Evamānayanāṅ cakre pūrvaṅ svīyaśiromaṇau, Bhāvanābhyāmatīspaṣṭarṅ saṁyaḡāryo’ pi Bhāskarāḡ*”  
—*SitVi*, ii. 70.

Or “This theorem, which is evident from the two *Bhāvanās*, was stated before also by the highly respected Bhāskara in his (*Siddhānta-*) *śiromaṇi*.”

<sup>4</sup>“*Tasya cānayanasyāryaiḡ siddhāntajñaiḡ puroditā, Vāsanaḡ bahubhiḡ svasvabuddhivaicitryataḡ sphuṭāḡ*” —*SitVi*, ii. 71.

Or “Many correct proofs of this theorem were given before by the learned authors of the *Siddhāntas* according to the manifoldness of their intelligence.”

<sup>5</sup>ii. 68-9 (Gloss).

$$\text{Now } PG^2 + QG^2 = QP^2 = QT^2 + PT^2.$$

Therefore, substituting the values

$$(kojy\bar{a} \beta - kojy\bar{a} \alpha)^2 + (jy\bar{a} \alpha + jy\bar{a} \beta)^2 = \{jy\bar{a} (\alpha + \beta)\}^2 + \{R - kojy\bar{a} (\alpha + \beta)\}^2$$

Simplifying we get

$$kojy\bar{a} (\alpha + \beta) = \frac{1}{R} (kojy\bar{a} \alpha \cdot kojy\bar{a} \beta - jy\bar{a} \alpha \cdot jy\bar{a} \beta)$$

which is equivalent to

$$\cos (\theta + \phi) = \cos \theta \cos \phi - \sin \theta \sin \phi.$$

Again

$$\begin{aligned} R^2 - \{kojy\bar{a} (\alpha + \beta)\}^2 &= \frac{1}{R^2} \{R^4 - (kojy\bar{a} \alpha \cdot kojy\bar{a} \beta - jy\bar{a} \alpha \cdot jy\bar{a} \beta)^2\} \\ &= \frac{1}{R^2} [ \{(jy\bar{a} \alpha)^2 + (kojy\bar{a} \alpha)^2\} \times \\ &\quad \{(jy\bar{a} \beta)^2 + (kojy\bar{a} \beta)^2\} - (kojy\bar{a} \alpha \cdot kojy\bar{a} \beta - jy\bar{a} \alpha \cdot jy\bar{a} \beta)^2 ] \end{aligned}$$

$$\text{or, } jy\bar{a} (\alpha + \beta) = \frac{1}{R} (jy\bar{a} \alpha \cdot kojy\bar{a} \beta + kojy\bar{a} \alpha \cdot jy\bar{a} \beta),$$

which is

$$\sin (\theta + \phi) = \sin \theta \cos \phi + \cos \theta \sin \phi.$$

Since

$$PG^2 + Q'G^2 = Q'P^2 = Q'T'^2 + PT'^2,$$

we have

$$(kojy\bar{a} \beta - kojy\bar{a} \alpha)^2 + (jy\bar{a} \alpha - jy\bar{a} \beta)^2 = \{jy\bar{a} (\alpha - \beta)\}^2 + \{R - kojy\bar{a} (\alpha - \beta)\}^2.$$

Therefore,

$$kojy\bar{a} (\alpha - \beta) = \frac{1}{R} (kojy\bar{a} \alpha \cdot kojy\bar{a} \beta + jy\bar{a} \alpha \cdot jy\bar{a} \beta),$$

which is

$$\cos (\theta - \phi) = \cos \theta \cos \phi + \sin \theta \sin \phi,$$

whence, proceeding as before, we get

$$jy\bar{a} (\alpha - \beta) = \frac{1}{R} (jy\bar{a} \alpha \cdot kojy\bar{a} \beta - kojy\bar{a} \alpha \cdot jy\bar{a} \beta)$$

or  $\sin (\theta - \phi) = \sin \theta \cos \phi - \cos \theta \sin \phi$ .

*Alternative Proof.* Let the arc  $YP = \text{arc } PP_1 = \beta$  and the arc  $YQ = \text{arc } QQ_1 = \alpha$ . Then it is obvious from the figure that

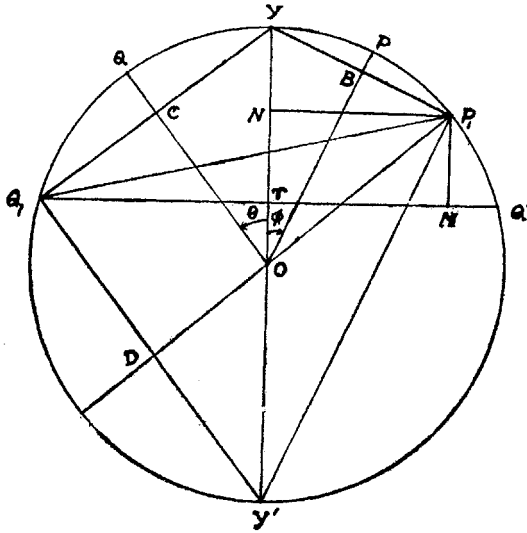


Fig. 8

$$YP_1 = 2 jy\bar{a} \beta, Y'P_1 = 2 kojy\bar{a} \beta,$$

$$YQ_1 = 2 jy\bar{a} \alpha, Y'Q_1 = 2 kojy\bar{a} \alpha.$$

Also

$$Q_1P_1^2 = Q_1D^2 + DP_1^2 = \{jy\bar{a} (2\alpha + 2\beta)\}^2 + \{(utjy\bar{a} (2\alpha + 2\beta))\}^2.$$

Therefore,  $Q_1P_1 = 2 jy\bar{a} (\alpha + \beta)$ .

Similarly  $Q'_1P_1 = 2 jy\bar{a} (\alpha - \beta)$ .

By the geometrical rules for finding the height and the segments of the base of a triangle whose sides are known, it can be easily proved that

$$Y'N = \frac{2}{R} (kojy\bar{a} \beta)^2, YN = \frac{2}{R} (jy\bar{a} \beta)^2, P_1N = \frac{2}{R} jy\bar{a} \beta \cdot kojy\bar{a} \beta,$$

$$Y'T = \frac{2}{R} (kojy\bar{a} \alpha)^2, YT = \frac{2}{R} (jy\bar{a} \alpha)^2, QT = \frac{2}{R} 2 jy\bar{a} \alpha \cdot kojy\bar{a} \alpha.$$

$$\begin{aligned}
\text{Now } Q_1P_1^2 &= Q_1M^2 + P_1M^2 \\
&= (Q_1T + P_1N)^2 + (YT - YN)^2 \\
&= \frac{4}{R^2} (jy\bar{a} \alpha. kojy\bar{a} \alpha + jy\bar{a} \beta. kojy\bar{a} \beta)^2 + \frac{4}{R^2} \{(jy\bar{a} \alpha)^2 - (jy\bar{a} \beta)^2\}^2 \\
&= \frac{4}{R^2} (jy\bar{a} \alpha. kojy\bar{a} \beta + kojy\bar{a} \alpha. jy\bar{a} \beta)^2
\end{aligned}$$

$$\therefore jy\bar{a} (\alpha + \beta) = \frac{1}{R} (jy\bar{a} \alpha. kojy\bar{a} \beta + kojy\bar{a} \alpha. jy\bar{a} \beta),$$

which is equivalent to

$$\sin (\theta + \phi) = \sin \theta \cos \phi + \cos \theta \sin \phi.$$

Again

$$\begin{aligned}
Q'P_1^2 &= Q'M^2 + MP_1^2 \\
&= (Q_1T - P_1N)^2 + (YT - YN)^2 \\
&= \frac{4}{R^2} (jy\bar{a} \alpha. kojy\bar{a} \alpha - jy\bar{a} \beta. kojy\bar{a} \beta)^2 + \frac{4}{R^2} \{(jy\bar{a} \alpha)^2 - (jy\bar{a} \beta)^2\}^2
\end{aligned}$$

whence

$$jy\bar{a} (\alpha - \beta) = \frac{1}{R} (jy\bar{a} \alpha. kojy\bar{a} \beta - kojy\bar{a} \alpha. jy\bar{a} \beta),$$

which is equivalent to

$$\sin (\theta - \phi) = \sin \theta \cos \phi - \cos \theta \sin \phi.$$

The above theorems are called *Bhāvanā* (“demonstration” or “proof” meaning “any thing demonstrated or proved”, hence “theorem”).

They are again divided into *Samāsa-bhāvanā* or *Yoga-bhāvanā* (“Addition Theorem”) and *Antara-bhāvanā* or *Viyoga-bhāvanā* (“Subtraction Theorem”).<sup>1</sup>

In the proofs given above the arcs  $\alpha$  and  $\beta$  have been tacitly assumed to be each less than  $90^\circ$ . But the theorems are quite general and hold true even when the arcs are greater than  $90^\circ$ .

Thus Kamalākara observes:

“Even when the two arcs go beyond  $90^\circ$  to any even or odd quadrant (the theorems) will remain the same, not otherwise. That is the opinion of those who are aware of the true facts.”<sup>2</sup>

<sup>1</sup>*SiŚi, Gola*, xiv. 21 (*Gloss*); *SiTVi*, ii. 65.

<sup>2</sup>*SiTVi*, ii. 66f.

*Functions of Multiple Angles*

As corollaries to the general case of the theorems for expanding  $\sin(\theta \pm \phi)$  and  $\cos(\theta \pm \phi)$ , Bhāskara II (1150) indicates how to derive the functions of multiple angles. He observes:

“This being proved, it becomes an argument for determining the values of other functions. For example, take the case of the combination of functions of equal arcs: by combining the functions of any arc with those of itself, we get the functions of twice that arc; by combining the functions of twice the arc with those of twice the arc, we get functions of four times that arc; and so on. Next take the case of combination of functions of unequal arcs: on combining the functions of twice an arc with those of thrice that arc, by the addition theorem we get the functions of five times that arc; but by the subtraction theorem, we get the functions of one time that arc; and so on.”<sup>1</sup>

The theorems meant here are clearly these:

$$\begin{aligned} \sin 2\theta &= 2 \sin \theta \cos \theta, \\ \cos 2\theta &= \cos^2 \theta - \sin^2 \theta, \\ \sin 4\theta &= 2 \sin 2\theta \cos 2\theta, \\ &= 4 \sin \theta \cos \theta (\cos^2 \theta - \sin^2 \theta), \\ \cos 4\theta &= \cos^2 2\theta - \sin^2 2\theta, \\ &= \cos^4 \theta - 6 \sin^2 \theta \cos^2 \theta + \sin^4 \theta, \\ \sin 3\theta &= 3 \sin \theta \cos^2 \theta - \sin^3 \theta, \\ \cos 3\theta &= \cos^3 \theta - 3 \cos \theta \sin^2 \theta, \\ \sin 5\theta &= \sin 2\theta \cos 3\theta + \cos 2\theta \sin 3\theta, \\ \cos 5\theta &= \cos 2\theta \cos 3\theta - \sin 2\theta \sin 3\theta, \\ \sin \theta &= \sin 3\theta \cos 2\theta - \cos 3\theta \sin 2\theta, \\ \cos \theta &= \cos 3\theta \cos 2\theta + \sin 3\theta \sin 2\theta. \end{aligned}$$

All these theorems have been expressly stated by Kamalākara (1658).

He says:

“Hereafter I shall describe how to find the *Rsine* of twice, thrice, four times or five times an arc, having known the *Rsine* of the sum of two arcs. The product of the *Rsine* and *Rcosine* of an arc is multiplied by 2 and divided by the radius; the result is the *Rsine* of twice that arc.”<sup>2</sup>

“The difference of the squares of the *Rsine* and *Rcosine* of an arc is divided by the radius; the quotient is certainly the *Rcosine* of twice that arc.”<sup>3</sup>

He has given the following proof of the above two formulae.<sup>4</sup>

<sup>1</sup>*SiŚi, Gola*, xiv. 21-2 (*Gloss*).

<sup>2</sup>*SiTVi*, ii. 73.

<sup>3</sup>*Ibid*, ii. 90.

<sup>4</sup>See his own gloss on the preceding rules.





By the successive application of the Addition Theorems, Kamalākara obtains the formulae:<sup>1</sup>

$$\begin{aligned} jyā\ 3\alpha &= \{3R^2\ jyā\ \alpha - 4(jyā\ \alpha)^3\}/R^2, \\ kojyā\ 3\alpha &= \{4\ (kojyā\ \alpha)^3 - 3R^2\ kojyā\ \alpha\}/R^2, \\ jyā\ 4\alpha &= 4\ \{(kojyā\ \alpha)^3\ jyā\ \alpha - (jyā\ \alpha)^3\ kojyā\ \alpha\}/R^3, \\ kojyā\ 4\alpha &= \{(kojyā\ \alpha)^4 - 6\ (kojyā\ \alpha)^2\ (jyā\ \alpha)^2 + (jyā\ \alpha)^4\}/R^3, \\ jyā\ 5\alpha &= \{(jyā\ \alpha)^5 - 10\ (jyā\ \alpha)^3\ (kojyā\ \alpha)^2 + 5\ jyā\ \alpha\ (kojyā\ \alpha)^4\}/R^4, \\ kojyā\ 5\alpha &= \{(kojyā\ \alpha)^5 - 10\ (kojyā\ \alpha)^3\ (jyā\ \alpha)^2 + 5\ kojyā\ \alpha\ (jyā\ \alpha)^4\}/R^4; \end{aligned}$$

which are of course equivalent to

$$\begin{aligned} \sin\ 3\theta &= 3\ \sin\ \theta - 4\ \sin^3\ \theta, \\ \cos\ 3\theta &= 4\ \cos^3\ \theta - 3\ \cos\ \theta, \\ \sin\ 4\theta &= 4\ (\cos^3\ \theta\ \sin\ \theta - \sin^3\ \theta\ \cos\ \theta), \\ \cos\ 4\theta &= \cos^4\ \theta - 6\ \cos^2\ \theta\ \sin^2\ \theta + \sin^4\ \theta, \\ \sin\ 5\theta &= \sin^5\ \theta - 10\ \sin^3\ \theta\ \cos^2\ \theta + 5\ \sin\ \theta\ \cos^4\ \theta, \\ \cos\ 5\theta &= \cos^5\ \theta - 10\ \cos^3\ \theta\ \sin^2\ \theta + 5\ \cos\ \theta\ \sin^4\ \theta. \end{aligned}$$

### Functions of Submultiple Angles

It has been stated before that the following two formulae for the sine of half an angle were known to almost all the Hindu astronomers:

$$\sin\ \frac{\theta}{2} = \frac{1}{2} \sqrt{\sin^2\ \theta + \text{versin}^2\ \theta},$$

$$\sin\ \frac{\theta}{2} = \sqrt{\frac{1}{2} (1 - \cos\ \theta)}.$$

Besides these<sup>2</sup> Kamalākara has given formulae for the functions of the third, fourth and fifth parts of an arc.

“Find the cube of one-third the Rsine of an arc; divide it by the square of the radius; the quotient is added to its one-third and the sum again to one-third the Rsine of the arc; the result is nearly the Rsine of one-third that arc. From the cube of this again further accurate values can be obtained.”<sup>3</sup>

$$jyā\ \frac{\alpha}{3} = \frac{1}{3} jyā\ \alpha + \frac{4}{3 R^2} \left( \frac{jyā\ \alpha}{3} \right)^3.$$

The rationale of this formula has been stated to be this: As has been proved before

$$jyā\ 3\beta = 3\ jyā\ \beta - \frac{4}{R^2} (jyā\ \beta)^3.$$

<sup>1</sup>SiTVi, ii. 75-7 and also the Gloss on them.

<sup>2</sup>SiTVi, ii. 78 f.

<sup>3</sup>Ibid, ii. 81.

Put  $3\beta = \alpha$ ; then this formula will become

$$jy\bar{a} \frac{\alpha}{3} = \frac{1}{3} jy\bar{a}\alpha + \frac{4}{3R^2} \left( jy\bar{a} \frac{\alpha}{3} \right)^3. \quad (i)$$

Now  $jy\bar{a} \frac{\alpha}{3}$  can be taken, says Kamalākara, as a *rough approximation* (*sthūla*) to be equal to  $\left( \frac{jy\bar{a} \alpha}{3} \right)^3$ . So that approximately

$$jy\bar{a} \frac{\alpha}{3} = \frac{1}{3} jy\bar{a} \alpha + \frac{4}{3R^2} \left( \frac{jy\bar{a} \alpha}{3} \right)^3, \quad (ii)$$

as stated in the rule. Very nearer approximation (*sūkṣmāsanna*) to the value of  $jy\bar{a} \frac{\alpha}{3}$  can be found by substituting the cube of this value in the last term of (i) and by repeating similar operations.

The form (ii) is equivalent to

$$\sin \frac{\theta}{3} = \frac{1}{3} \sin \theta + \frac{4}{81} \sin^3 \theta.$$

“From the known value of the Rsine of an arc, first calculate the value of the Rsine of half that arc; the Rsine of the arc is divided by that and multiplied by the square of the radius; the result is subtracted from twice the square of the radius. Half the square-root of the remainder is the value of the Rsine of one-fourth that arc.”<sup>1</sup>

$$jy\bar{a} \frac{\alpha}{4} = \frac{1}{2} \sqrt{2R^2 - R^2 \frac{jy\bar{a} \alpha}{jy\bar{a} (\alpha/2)}}.$$

The *rationale* of this formula is given thus: It is known that

$$\begin{aligned} jy\bar{a} 4\beta &= \frac{4}{R^3} \{ (kojy\bar{a} \beta)^3 jy\bar{a} \beta - (jy\bar{a} \beta)^3 kojy\bar{a} \beta \}, \\ &= \frac{4}{R^3} \{ R^2 jy\bar{a} \beta kojy\bar{a} \beta - 2 (jy\bar{a} \beta)^3 kojy\bar{a} \beta \}, \end{aligned}$$

Putting  $\alpha$  for  $4\beta$ , we get

$$\begin{aligned} R^3 jy\bar{a} \alpha &= 4 jy\bar{a} \frac{\alpha}{4} kojy\bar{a} \frac{\alpha}{4} \{ R^2 - 2 (jy\bar{a} \frac{\alpha}{4})^2 \}, \\ &= 2R jy\bar{a} \frac{\alpha}{2} \{ R^2 - 2 (jy\bar{a} \frac{\alpha}{4})^2 \}; \end{aligned}$$

<sup>1</sup>SiTVi, ii. 82-83.

whence

$$jy\bar{a} \frac{\alpha}{4} = \frac{1}{2} \sqrt{2R^2 - R^2 (jy\bar{a} \alpha) / jy\bar{a} \frac{\alpha}{2}}$$

or

$$\sin \frac{\theta}{4} = \frac{1}{2} \sqrt{2 - \frac{\sin \theta}{\sin (\theta/2)}}$$

“The intelligent should first find the one-fifth of the *Rsine* of the given arc; divide four times the cube of that by the square of the radius; the quotient should be called the “first”. Multiply the “first” by the square of the fifth part of the *Rsine* and divide the product by the square of the radius; lessen this quotient by its fifth part and mark the remainder as the “second”. One-fifth of the *Rsine* of the arc added with the “first” and diminished by the “second”, will be clearly the value of the *Rsine* of the fifth part of the arc. Finding the value of the “first” again from this, further approximate value to the *R sine* of one-fifth the arc can be found. Still closer approximations can be obtained by repeating the process stated above.”<sup>1</sup>

$$jy\bar{a} \frac{\alpha}{5} = \frac{1}{5} jy\bar{a} \alpha + \frac{4}{R^2} \left( \frac{jy\bar{a} \alpha}{5} \right)^3 - \frac{16}{5R^4} \left( \frac{jy\bar{a} \alpha}{5} \right)^5.$$

The *rationale* is stated to be this: It has been established before that  $R^4 jy\bar{a} 5\beta = (jy\bar{a} \beta)^5 - 10 (jy\bar{a} \beta)^3 (kojy\bar{a} \beta)^2 + 5jy\bar{a} \beta (kojy\bar{a} \beta)^4$ .

Substituting the value  $R^2 - (jy\bar{a} \beta)^2$  for  $(kojy\bar{a} \beta)^2$  in this, we get  $R^4 jy\bar{a} 5\beta = 16 (jy\bar{a} \beta)^5 - 20R^2 (jy\bar{a} \beta)^3 + 5R^4 jy\bar{a} \beta$

Putting  $\alpha$  for  $5\beta$ ,

$$jy\bar{a} \frac{\alpha}{5} = \frac{1}{5} jy\bar{a} \alpha + \frac{4}{R^2} (jy\bar{a} \frac{\alpha}{5})^3 - \frac{61}{5R^4} (jy\bar{a} \frac{\alpha}{5})^5. \tag{i}$$

In the last two terms on the right hand side, one may take as a rough approximation

$$jy\bar{a} \frac{\alpha}{5} = \frac{1}{5} jy\bar{a} \alpha;$$

so that

$$jy\bar{a} \frac{\alpha}{5} = \frac{1}{5} jy\bar{a} \alpha + \frac{4}{R^2} \left( \frac{jy\bar{a} \alpha}{5} \right)^3 - \frac{16}{5R^4} \left( \frac{jy\bar{a} \alpha}{5} \right)^5. \tag{ii}$$

---

<sup>1</sup>*Sitvi*, ii. 84-87.

Again substituting this value of  $jy\bar{a} \frac{\alpha}{5}$  in the last two terms of (i) and repeating similar operations, closer approximations to the value of  $jy\bar{a} \frac{\alpha}{5}$  can be obtained.

The formula (ii) is equivalent to

$$\sin \frac{\theta}{5} = \frac{1}{5} \sin \theta + 4 \left( \frac{\sin \theta}{5} \right)^3 - \frac{16}{5} \left( \frac{\sin \theta}{5} \right)^5.$$

Kamalākara then observes that “in this way, the Rsines of other desired submultiples of an arc should be obtained.”<sup>1</sup>

$$\text{Sin } \frac{(\theta - \phi)}{2}.$$

*Bhāskara II says:*

“Find the difference of the Rsines of two arcs and also of their Rcosines; then find the square-root of the sum of the squares of the two results; half this root will be the Rsine of half the difference of the two arcs.”<sup>2</sup>

That is,

$$jy\bar{a} \frac{1}{2} (\alpha - \beta) = \frac{1}{2} \left\{ (jy\bar{a} \alpha - jy\bar{a} \beta)^2 + (kojy\bar{a} \alpha - kojy\bar{a} \beta)^2 \right\}^{\frac{1}{2}}.$$

or, in modern notations,

$$\sin \frac{1}{2}(\theta - \phi) = \frac{1}{2} \left\{ (\sin \theta - \sin \phi)^2 + (\cos \theta - \cos \phi)^2 \right\}^{\frac{1}{2}}.$$

Kamalākara writes:

“Half the square-root of the sum of the squares of the differences of Rsines and Rcosines of two arcs is certainly equal to the Rsine of half the difference of the two arcs.”<sup>3</sup>

The latter has given the following proof of it.<sup>4</sup> Let the arc  $XP$  be denoted by  $\beta$  and the arc  $XQ$  by  $\alpha$ ; then

<sup>1</sup>*SitVi*, ii, 87 (c-d).

<sup>2</sup>*SiSi, Gola*, xiv, 13.

<sup>3</sup>*SitVi*, ii, 94.

<sup>4</sup>*Ibid*, (Gloss).

$$QC = QT - PM = jy\bar{a} \alpha - jy\bar{a} \beta,$$

$$PC = OM - OT = kojy\bar{a} \beta - kojy\bar{a} \alpha.$$

Now,

$$PQ^2 = QC^2 + PC^2.$$

Hence,

$$jy\bar{a} \frac{1}{2} (\alpha - \beta) = \frac{1}{2} \{ (jy\bar{a} \alpha - jy\bar{a} \beta)^2 + (kojy\bar{a} \alpha - kojy\bar{a} \beta)^2 \}^{\frac{1}{2}},$$

which is equivalent to

$$\sin \frac{1}{2} (\theta - \phi) = \frac{1}{2} \{ (\sin \theta - \sin \phi)^2 + (\cos \theta - \cos \phi)^2 \}^{\frac{1}{2}}.$$

*Theorem of Sines*

Brahmagupta<sup>1</sup> has made use of the important relation

$$\frac{a}{jy\bar{a} A} = \frac{b}{jy\bar{a} B} = \frac{c}{jy\bar{a} C}$$

which is of course equivalent to

$$\frac{a}{\sin A} = \frac{b}{\sin B} = \frac{c}{\sin C}$$

between the sides ( $a, b, c$ ) and angles ( $A, B, C$ ) of a plane triangle.

#### 4. FUNCTIONS OF PARTICULAR ANGLES

*Sine of 30°, 45° and 60°.* Preliminary to the calculation of tables of trigonometrical functions almost all the Hindu writers have stated the values of the Rsines of 30°, 45° and 60°.

$$jy\bar{a} 30^\circ = \sqrt{R^2/4}, \quad jy\bar{a} 45^\circ = \sqrt{R^2/2}, \quad jy\bar{a} 60^\circ = \sqrt{3R^2/4} \text{ or, in modern notations,}$$

$$\sin 30^\circ = \frac{1}{2}, \quad \sin 45^\circ = 1/\sqrt{2}, \quad \sin 60^\circ = \frac{1}{2}\sqrt{3}.$$

Śrīpati indicates the proof thus:

“The experts in spherics say that the circum-radius of a regular hexagon is equal to a side. So it will be perceived that the chord of the sixth part of the circumference

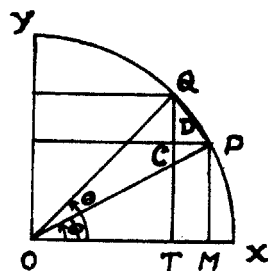


Fig.10

<sup>1</sup>KK, Part I, viii. 2. Our attention to this was first drawn by Professor P. C. Sen Gupta, who was then preparing a new edition of *Khandakhadyaka* with English translation and critical notes. This rule occurs in the works of other Hindu astronomers also.

of a circle is equal to its semi-diameter. The hypotenuse arising from the base and perpendicular (of a right-angled triangle) each equal to the semi-diameter is the chord of the fourth part of the circumference. Half those (chords) will be the *Rsines* of half those arcs."<sup>1</sup>

The same proof is also given by Bhāskara II:<sup>2</sup>

"The side of a regular hexagon inscribed in a circle is equal to its radius; this is well known and has also been stated in (my) Arithmetic. Hence follows that the *Rsine* of 30° is half the radius."

"Suppose a right-angled triangle whose base and perpendicular are each equal to the radius; the square-root of the sum of the squares of these will be equal to the side of a square inscribed in that circle and it is again the chord of 90°. Take the half of that. Hence the sum of the squares (of the sides) is divided by four; and the result is half the square of the radius. The square-root of that, it thus follows, is the *Rsine* of 45°."

"The *Rsine* of 60° is equal to the *Rcosine* of 30°, the *Rsine* of which is equal to half the semi-diameter."

*Sin 18° and sin 36°*. Bhāskara II says:

"The square-root of five times the square of the radius is diminished by the radius and the remainder is divided by four; the result is the exact value of the *Rsine* of 18°."<sup>3</sup>

$$jyā\ 18^\circ = \frac{1}{4} (\sqrt{5R^2} - R);$$

or  $\sin 18^\circ = \frac{1}{4}(\sqrt{5} - 1).$

"The square-root of five times the square of the square of the radius is subtracted from five times the square of the radius and the remainder is divided by eight; the square-root of the quotient is the *R sine* of 36°."

"Or the radius multiplied by 5878 and divided by 10000, is the *Rsine* of 36°. The *Rcosine* of that is the *Rsine* of 54°."<sup>4</sup>

$$jyā\ 36^\circ = \sqrt{\frac{1}{8} (5R^2 - \sqrt{5R^4})} = \frac{5878R}{10000}$$

$$\text{That is, } \sin 36^\circ = \sqrt{\frac{1}{8} (5 - \sqrt{5})} = \frac{5878}{10000}$$

<sup>1</sup>*SiSe*, xvi. 11-2.

<sup>2</sup>*SiSi, Gola*, v. 3-4 (Gloss).

<sup>3</sup>*Ibid, Gola*, xiv. 9.

<sup>4</sup>*Ibid, Gola*, xiv. 7-8.

Since  $\sqrt{5} = 2.237411$  approximately

$$\therefore 5 - \sqrt{5} = 2.762589 \dots$$

$$\therefore \sqrt{\frac{1}{8}(5 - \sqrt{5})} = \sqrt{.345323} \dots = .5878 \text{ approximately.}$$

Kamalākara proved the results thus:

Let  $x$  denote  $jyā 18^\circ$ ; then

$$\frac{1}{2} R (R - x) = \frac{1}{2} R. utjyā 72^\circ = (jyā 36^\circ)^2;$$

$$\frac{2x^2}{R} = R - kojyā 36^\circ = utjyā 36^\circ;$$

$$\begin{aligned} \frac{1}{2} R (R - x) + \left( \frac{2x^2}{R} \right)^2 &= (jyā 36^\circ)^2 + (utjyā 36^\circ)^2 \\ &= 4 (jyā 18^\circ)^2 \\ &= 4x^2, \end{aligned}$$

$$\text{or } 8 x^2 R^2 = 8 x^4 - R^3 x + R^4,$$

or, multiplying by 8 and arranging,

$$16R^2 x^2 + 8 R^3 x + R^4 = 9 R^4 - 48 R^2 x^2 + 64 x^4,$$

whence taking the positive square roots of the two sides, we get

$$\begin{aligned} 4 R x + R^2 &= 3 R^2 - 8 x^2, \\ \text{or } (4 x + R)^2 &= 5 R^2. \end{aligned}$$

$$\text{Therefore } x = \frac{1}{4} (\sqrt{5R^2} - R);$$

the other sign is neglected since  $x$  must be less than  $R$ .

$$\text{Again } (jyā 36^\circ)^2 = \frac{1}{2} R(R - x),$$

$$\begin{aligned} &= \frac{R}{8} (5R - \sqrt{5R^2}), \\ \therefore jyā 36^\circ &= \sqrt{\frac{1}{8}(5R^2 - \sqrt{5R^4})} \end{aligned}$$

*Sin  $\pi/N$*

In his treatise on arithmetic, Bhāskara II has given a rule which yields the Rsine of certain particular angles to a very fair degree of approximation.

“Multiply the diameter of a circle by 103923, 84853, 70534, 60000, 52055, 45922 and 41031 severally and divide the products by 120000; the quotients will be the sides of regular polygons inscribed in the circle from the triangle to the enneagon respectively.”<sup>1</sup>

If  $S_n$  be a side of a regular polygon of  $n$  sides inscribed in a circle of diameter  $D$ , then according to Bhāskara II,

$$S_3 = D \frac{103923}{120000} = D \times .866025$$

$$S_4 = D \frac{84853}{120000} = D \times .707108\dot{8}$$

$$S_5 = D \frac{70534}{120000} = D \times .58778\dot{8}$$

$$S_6 = D \frac{60000}{120000} = D \times .5$$

$$S_7 = D \frac{52055}{120000} = D \times .433791\dot{6}$$

$$S_8 = D \frac{45922}{120000} = D \times .38268\dot{8}$$

$$S_9 = D \frac{41031}{120000} = D \times .341925$$

where are given the formulae of Bhāskara II first in their original forms and then in decimals. Now, we know that

$$s_n = D \sin \frac{\pi}{n}.$$

Hence it is found that

$$\sin 60^\circ = .866025$$

$$\sin 45^\circ = .707108\dot{8}$$

$$\sin 36^\circ = .58778\dot{8}$$

$$\sin \pi/7 = .433791\dot{6}$$

$$\sin \pi/8 = .38268\dot{8}$$

$$\sin \pi/9 = .341925$$

According to modern computation

$$\sin 60^\circ = .8660254\dots$$

$$\sin 45^\circ = .7071067\dots$$

$$\sin 36^\circ = .5877853$$

$$\sin \pi/7 = .4338819$$

$$\sin \pi/8 = .3826834$$

$$\sin \pi/9 = .3420201$$

<sup>1</sup> L, vss. 206-7, p. 207.



Comparing the two tables we find that except in case of  $\sin \pi/7$  and  $\sin \pi/9$  Bhāskara's approximations are correct up to five places of decimals; in these two latter cases the results are near enough.

### *Approximate Formula of Bhāskara I*

Bhāskara I (629) has given the following rule for the calculation of the *Rsine* and *Rcosine* of an arc without the help of a table.

"Subtract the arc in degrees from the degrees of the semi-circumference and multiplying the arc by the remainder, put down (the result) at two places, (at one place) subtract (the quantity) from 40500; by one-fourth of the remainder divide the quantity (at the second place) multiplied by the maximum value of the function; thus the value of the direct or reversed *Rsine* of an arc and its complement is obtained wholly."<sup>1</sup>

If  $\alpha$  be an arc of a circle of radius  $R$  in terms of degrees, then

$$jyā \alpha = \frac{R(C/2 - \alpha)\alpha}{\{40500 - (C/2 - \alpha)\alpha\}/4}$$

where  $C$  denotes the circumference of the circle in terms of degrees. Since  $40500 = (5/4) \times 180 \times 180$ , we can write the formula in the form

$$jyā \alpha = \frac{4R(C/2 - \alpha)\alpha}{5/4 (C/2)^2 - (C/2 - \alpha)\alpha}$$

which is of course equivalent to

$$\sin \theta = \frac{4(\pi - \theta)\theta}{(5/4)\pi^2 - (\pi - \theta)\theta}$$

From a statement of Bhāskara I it appears that this formula was known to Āryabhaṭa I.<sup>2</sup>

The above formula has been restated by Brahmagupta (628) thus:

"Subtract the degrees of an arc or its complement from the semicircle (i.e. 180) and multiply (the remainder) by that; subtract one-fourth the product from 10125; divide the product by the remainder and multiply by the semi-diameter; (the result) is the *Rsine* of that (arc or its complement)."<sup>3</sup>

$$jyā \alpha = \frac{R(180 - \alpha)\alpha}{10125 - (180 - \alpha)\alpha/4}$$

<sup>1</sup>*MBh*, vii. 17ff.

<sup>2</sup>Bhāskara I's com. on *A*, i. 11, p. 40.

<sup>3</sup>*BrSpŚi*, xiv. 23.

Almost in the same way Śrīpati (1039) says:

“Subtract the degrees of an arc or its complement from 180 and multiply (the remainder) by that; subtract one-fourth the product from 10125; multiply the product by the semi-diameter and divide by this remainder; thus the *R*sine of an arc or its complement can be found even without (a table of *R*sines).”<sup>1</sup>

Bhāskara II (1150) writes:

“Subtract an arc from the circumference and multiply (the remainder) by the arc; this product is called the ‘first’. From five times the fourth part of the square of the circumference subtract the ‘first’, and by the remainder divide the ‘first’ multiplied by four times the diameter; the quotient will be the chord of the arc.”<sup>2</sup>

If  $s$  denote the chord of an arc  $\beta$  of a circle, then

$$s = \frac{8R(C-\beta)\beta}{\frac{5}{4}C^2 - (C-\beta)\beta}$$

Now if  $\beta=2\alpha$ , then  $s=2jy\bar{a}\alpha$ . So that on making the substitutions this formula will easily reduce to that of the elder Bhāskara.

This formula has been used by Gaṇeśa (1545) in his *Grahalāghava*.<sup>3</sup> Though it gives only a roughly approximate (*sthūla*) value of the *R*sine of an arc, observes Bhāskara II, it simplifies operations.

On putting  $\theta = \pi/2 - \phi$ , in the above approximate formula, it becomes

$$\begin{aligned} \cos \phi &= \frac{16(\pi/2 + \phi)(\pi/2 - \phi)}{5\pi^2 - 4(\pi/2 + \phi)(\pi/2 - \phi)} \\ &= \frac{\pi^2 - 4\phi^2}{\pi^2 + \phi^2} \\ &= \left(1 - \frac{4\phi^2}{\pi^2}\right) \left(1 - \frac{\phi^2}{\pi^2} + \frac{\phi^4}{\pi^4}\right), \end{aligned}$$

neglecting higher powers. Therefore, to the same order of approximation,

$$\cos \phi = 1 - \frac{5\phi^2}{\pi^2} + \frac{5\phi^4}{\pi^4}.$$

<sup>1</sup>*SiŚe*, iii. 17.

<sup>2</sup>*L*, vs. 210, p. 21e. Also see *GK*, par 2, pp. 80-81.

<sup>3</sup>*GrL*, ii. 2 f.

If we put  $\pi = \sqrt{10}$  approximately, we get

$$\cos \phi = 1 - \frac{\phi^2}{2} + \frac{\phi^4}{20}.$$

nearly. According to modern Trigonometry, to the same order of approximation,

$$\cos \phi = 1 - \frac{\phi^2}{2} + \frac{\phi^4}{24}.$$

Again putting  $\phi = \pi/n$  in Bhāskara I's formula, where  $n$  is an integer, we get

$$\sin \pi/n = \frac{16(n-1)}{5n^2 - 4(n-1)}$$

whence we have

$\sin \pi/7 = \cdot 4343 \dots$ ,  $\sin \pi/8 = \cdot 3835 \dots$ ,  $\sin \pi/9 = \cdot 3431 \dots$ ,  
which are correct up to two places of decimals, the third figure in every case being too large.

#### *Inverse Formula of Brahmagupta*

Brahmagupta gave the following rule for finding approximately the arc corresponding to a given Rsine function:

“Multiply 10125 by the given Rsine and divide by the quarter of the given Rsine plus the radius; subtracting the quotient from the square of 90, extract the square-root and subtract (the root) from 90; the remainder will be in degrees and minutes; thus will be found the arc of the given Rsine without the table of Rsines.”<sup>1</sup>

If  $\alpha$  be the arc corresponding to the given Rsine function  $m$ , then the rule says that

$$\alpha = 90 - \sqrt{8100 - \frac{10125m}{(m/4+r)}}.$$

This result follows easily on reversing the approximate formula for the Rsine and was very likely obtained in the same way.

$$m = jyā \alpha = \frac{R(180 - \alpha) \alpha}{10125 - (180 - \alpha) \alpha / 4}.$$

Then,

$$\alpha^2 - 180\alpha + \frac{10125m}{(m/4+r)} = 0.$$

<sup>1</sup>*BrSpSi*, xiv. 25-6.

Therefore,

$$\alpha = 90 - \sqrt{8100 - \frac{10125m}{(m/4+r)}}.$$

The negative sign of the radical being retained, since  $\alpha$  is supposed to be less than  $90^\circ$ .

Śrīpati describes the inverse formula thus:

“Multiply 10125 by the given *Rsine* and divide by the quarter of the given *Rsine* plus the radius; then subtract the quotient from the square of 90; ninety degrees lessened by the square root (of the remainder) will be the arc (determined) without the table of *Rsines*.”<sup>1</sup>

Bhāskara II writes:

“By four times the diameter added with the chord divide the square of the circumference multiplied by five times a quarter of the chord; the quotient being subtracted from the fourth part of the square of the circumference, and the square-root of the remainder being diminished from half the circumference, the result will be the arc.”<sup>2</sup>

That is:

$$\beta = \frac{C}{2} - \left\{ \frac{C^2}{4} - \frac{5s C^2}{4(8R+s)} \right\}^{\frac{1}{2}}$$

which follows at once from his form of the approximate formula for the chord  $s$ .

## 5. TRIGONOMETRICAL TABLES

### *Twenty-four Sines*

The Hindus generally calculate tables of trigonometrical functions for every arc of  $3^\circ 45'$ , or what they call twenty-four *Rsines* in a quadrant. In the choice of 24, they seem to have been led by an ancient observation that “the ninety-sixth part of a circle looks (straight) like a rod”. Thus Balabhadra (c. 700 A.D.) observes, “If anybody asks the reason of this, he must know that each of these *Kardajat* is  $1/96$  of the circle = 225 minutes (=  $3\frac{3}{4}$  degrees). And if we reckon its *Rsine*, we find it also to be 225 minutes.”<sup>3</sup>

The origin of this idea again lies in the impression that the human eye-sight reaches to a distance of  $1/96$ th part of the circumference of the earth which appears

<sup>1</sup>SiSe, iii. 18.

<sup>2</sup>L, vs. 212, p. 216.

<sup>3</sup>Quoted by Al-Bīrūnī in his *India* (Sachau, *Alberuni's India*, I, p. 275). Balabhadra's works are now lost. According to Chambers' *Mathematical Tables*, we find  $\sin(3^\circ 45') = \cdot 0654031$ ,  $\tan(3^\circ 45') = \cdot 0655435$ ,  $\text{radian}(3^\circ 45') = \cdot 0654498$ , so that the assumption is fairly accurate.



Or,

$$\begin{aligned} SL &= NS - OS \left( \frac{QP}{OQ} \right)^2 \\ &= (QT - PM) - OS \left( \frac{QP}{OQ} \right)^2 \end{aligned}$$

Therefore,

$$RK = QT + (QT - PM) - QT \left( \frac{QP}{OQ} \right)^2$$

Now suppose the arc  $XQ = n\alpha$ ; then arc  $XP = (n-1)\alpha$ ;  $XR = (n+1)\alpha$ ;

further  $QP = 2jy\bar{a} \frac{\alpha}{2}$ . Hence

$$jy\bar{a} (n+1)\alpha = jy\bar{a} n\alpha + \{ jy\bar{a} n\alpha - jy\bar{a}(n-1)\alpha \} - jy\bar{a} n\alpha \times \left( \frac{2jy\bar{a} \alpha/2}{R} \right)^2,$$

which is equivalent to

$$\sin (n+1) \theta = \sin n\theta + \{ \sin n\theta - \sin (n-1)\theta \} - \sin n\theta \left( 2 \sin \frac{\theta}{2} \right)^2.$$

It is also probable that the formula was obtained trigonometrically thus:

$$jy\bar{a}(\xi \pm \eta) = \frac{1}{R} (jy\bar{a} \xi \text{ kojy}\bar{a} \eta \pm \text{kojy}\bar{a} \xi jy\bar{a} \eta).$$

Then,

$$jy\bar{a} (\xi + \eta) - jy\bar{a} \xi = \frac{1}{R} (\text{kojy}\bar{a} \xi jy\bar{a} \eta - jy\bar{a} \xi \text{ utjy}\bar{a} \eta),$$

and,

$$jy\bar{a} \xi - jy\bar{a} (\xi - \eta) = \frac{1}{R} (\text{kojy}\bar{a} \xi jy\bar{a} \eta + jy\bar{a} \xi \text{ utjy}\bar{a} \eta).$$

Hence;

$$\begin{aligned} jy\bar{a} (\xi + \eta) - jy\bar{a} \xi &= jy\bar{a} \xi - jy\bar{a} (\xi - \eta) - \frac{2 jy\bar{a} \xi \text{ utjy}\bar{a} \eta}{R} \\ &= jy\bar{a} \xi - jy\bar{a} (\xi - \eta) - jy\bar{a} \xi \left( \frac{2 jy\bar{a} \eta}{R} \right)^2 \end{aligned}$$

Now put  $\eta = \alpha$ ,  $\xi = n\alpha$ ; so that the formula becomes

$$jy\bar{a} (n+1) \alpha - jy\bar{a} n \alpha = jy\bar{a} n \alpha - jy\bar{a} (n-1) \alpha - jy\bar{a} n \alpha \left( \frac{2 jy\bar{a} \alpha}{R} \right)^2$$

So far the formula is mathematically accurate. According to the *Sūrya-siddhānta*

$$\alpha = 3^\circ 45' = 225', \text{ jyā } \alpha = 225', R = 3438'$$

Therefore

$$\begin{aligned} \left( \frac{2 \text{ jyā } \alpha / 2}{R} \right)^2 &= \left( \frac{\text{ jyā } \alpha}{R} \right)^2 \text{ approximately} \\ &= \left( \frac{225}{3438} \right)^2 = \left( \frac{1}{15.28} \right)^2 = \frac{1}{225} \text{ approximately.} \end{aligned}$$

Hence we get

$$\sin (n + 1) \theta = \sin n \theta + \left\{ \sin n \theta - \sin (n - 1) \theta \right\} - \frac{\sin n \theta}{225},$$

where

$$\theta = 3^\circ 45' \text{ and } n = 1, 2, \dots, 24.$$

According to modern calculation, the divisor in the last term will be slightly different. For

$$\left( 2 \sin \frac{\theta}{2} \right)^2 = (2 \sin 1^\circ 52' 30'')^2 = .00428255 = \frac{1}{233.506}, \text{ nearly.}$$

This little discrepancy, however, does not make much difference in the values of the *R*sine functions calculated on the basis of that formula. They are indeed fairly accurate even according to modern calculations except in a few instances.<sup>1</sup>

About this method of constructing the tables of *R* sines, Delambre remarks: "The method is curious; it indicates a method of calculating the table of sines by means of their second differences."<sup>2</sup> He then goes on: "This differential process has not up to now been employed except by Briggs who himself did not know that the constant factor was the square of the chord  $\triangle A$  ( $=3^\circ 45'$ ) or of the interval, and who could not obtain it except by comparing the second differences obtained in a different manner. The Indians also have probably done the same; they obtained the method of differences only from a table calculated previously by a geometric process. Here then is a method which the Indians possessed and which is found neither amongst the Greeks, nor amongst the Arabs."<sup>3</sup>

We do not understand what valid grounds were there for Delambre to suppose that the Hindus discovered the above theorem of *R*sines by inspection after having calculated the table of *R*sines by a different method. For there is absolutely no doubt

<sup>1</sup>*Vide infra.*

<sup>2</sup>Delambre, *Histoire de l'Astronomie Ancienne*, t. 1. Paris, 1817, p. 457.

<sup>3</sup>*Ibid.*, p. 459 f.

that the ancient Hindus were in possession of necessary and sufficient equipments to derive it in either of the ways indicated above. It is noteworthy that that theorem has an important geometrical foundation. If there be three arcs of a circle in arithmetical progression the sum of the sines of the two extreme arcs is to the sine of the middle arc as the sine of twice the common difference is to the sine of that difference. For

$$\begin{aligned} jy\bar{a} (\xi + \eta) + jy\bar{a} (\xi - \eta) &\equiv 2 jy\bar{a} \xi - \frac{2 jy\bar{a} \xi \sin jy\bar{a} \eta}{R} \\ &= \frac{2 jy\bar{a} \xi \cos jy\bar{a} \eta}{R} \end{aligned}$$

Hence,

$$\frac{jy\bar{a} (\xi + \eta) + jy\bar{a} (\xi - \eta)}{jy\bar{a} \xi} = \frac{2 \cos jy\bar{a} \eta}{R} = \frac{jy\bar{a} 2 \eta}{jy\bar{a} \eta}$$

This very remarkable property of the circle was discovered in Europe by Vieta (1580)<sup>1</sup>.

### Āryabhaṭa

The trigonometrical table of Āryabhaṭa I (499) contains the differences between the successive *R*sines for arcs of every 3° 45' of a circle of radius 3438'.<sup>2</sup> His first method of computing it, which is rather cryptic, seems to be the same as that followed by Varāhamihira (*infra*). The other is practically the same as that of the *Sūrya-siddhānta*, though put in a different form. He says:

"Divide a quarter of the circumference of a plane circle (into as many equal parts as desired). From (right) triangles and quadrilaterals (can be obtained) the *R*sines of equal arcs, as many as desired, for (any given) radius."<sup>3</sup>

What is meant by the author is very probably this: If *P* be any point on the arc of the quadrant, draw the perpendiculars *PM* and *PN*; also join *PX*. So that corresponding to *P* we have a rectangle *PMON* and a right-angled triangle *PMX*. Now having given the *R*sine (*PM*) of the arc *XP* (=  $\alpha$ ), we can determine from the rectangle *PMON* the side *PN* which is the sine of the arc ( $90^\circ - \alpha$ ). Having found *PN*, we can calculate *MX*, which is equal to  $R - jy\bar{a} (90^\circ - \alpha)$ . Then in the right-angled triangle *PMX*, we can determine the chord *PX*. Half of this is  $jy\bar{a} \alpha/2$ . Again

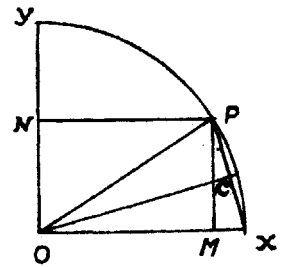


Fig. 12

<sup>1</sup>Playfair, J "Observations on the Trigonometrical Tables of the Brahmins", *Trans. Roy. Soc. Edin.*, iv (1798), pp. 83-106; compare also *Asiatic Researches*, iv. p. 165.

<sup>2</sup>*A.*, i. 12.

<sup>3</sup>*Ibid.*, ii. 11.



from a similar set of a rectangle and a right-angled triangle corresponding to the half arc, we can calculate  $jyā (90^\circ - \alpha/2)$  and  $jyā \alpha/4$ . Proceeding thus we can compute the Rsines of as many equal arcs as we please and it is clear that in so doing the quadrant will be broken up into a system of right-angled triangles and rectangles, as contemplated in the rule.

This is the interpretation of Āryabhaṭa's rule by his ancient commentators, like Sūryadeva Yajvā and Parameśvara (1430). Another interpretation will be this: The quadrant is trisected by the inscribed equilateral triangle and bisected by the inscribed square. The length of the arc between these points is  $15^\circ (=45^\circ - 30^\circ)$ . One-fourth of this is  $3^\circ 45'$ . So that the rule under discussion indicates how to divide the quarter of the circumference into portions of  $3^\circ 45'$  each. If this interpretation is right<sup>1</sup>, which is rather forced, then it will have to be said that Āryabhaṭa I gave only one method of computing the trigonometrical table.<sup>2</sup>

The *second* method of Āryabhaṭa I is this:

“The first Rsine divided by itself and then diminished by the quotient will give the second difference (of tabular Rsines). For computing any other difference, (the sum of) all the preceding differences is divided by the first Rsine and the quotient is subtracted from the preceding difference. Thus, all the remaining differences (can be calculated).”<sup>3</sup>

Let  $\Delta_1, \Delta_2, \dots, \Delta_n$  denote successive differences of the tabular Rsines, such that,  $\alpha$  being equal to  $3^\circ 45'$ ,

$$\begin{aligned} \Delta_1 &= jyā \alpha - jyā 0, \\ \Delta_2 &= jyā 2\alpha - jyā \alpha, \\ &\dots\dots\dots \\ \Delta_n &= jyā n\alpha - jyā (n-1)\alpha \end{aligned}$$

Then  $jyā n\alpha = \Delta_1 + \Delta_2 + \dots + \Delta_n$ .

The rule says:

$$\Delta_{n+1} = \Delta_n - \frac{\Delta_1 + \Delta_2 + \dots + \Delta_n}{jyā \alpha}$$

On substituting the values, this formula will be found to be equivalent to

$$\{jyā (n+1)\alpha - jyā n\alpha\} = \{jyā n\alpha - jyā (n-1)\alpha\} - \frac{jyā n\alpha}{jyā \alpha}$$

<sup>1</sup>This interpretation has been suggested by Rodet, Kaye and Sengupta.

<sup>2</sup>In this connection, the reader is referred to “Āryabhaṭīya of Āryabhaṭa,” edited with English translation by K. S. Shukla and K. V. Sarma, INSA, New Delhi, 1976, pp. 45-51.

<sup>3</sup>A, ii. 12.

It is also noteworthy that the text also admits of the following interpretation:

“The first *Rsine* is divided by itself and then diminished by the quotient; the result with the first *Rsine* will give the second *Rsine*. For (computing), any of the remaining *Rsines*, the sum of all the *Rsines* preceding it is divided by the first *Rsine* and the quotient is subtracted from the first *Rsine*, and the result added to the preceding *Rsine*.”

$$jy\bar{a} (n+1)\alpha = jy\bar{a} n\alpha + jy\bar{a} \alpha - (jy\bar{a} \alpha + jy\bar{a} 2\alpha + \dots + jy\bar{a} n\alpha) / jy\bar{a} \alpha.$$

If  $\Delta_1, \Delta_2, \dots$  be the tabular differences as before, then

$$\Delta_1 - \Delta_2 = \frac{2jy\bar{a} \alpha (R - kojy\bar{a} \alpha)}{R}$$

$$\Delta_2 - \Delta_3 = \frac{2jy\bar{a} 2\alpha (R - kojy\bar{a} \alpha)}{R}$$

.....

$$\Delta_n - \Delta_{n+1} = 2jy\bar{a} n\alpha \frac{R - kojy\bar{a} \alpha}{R}$$

Adding up, we get

$$\Delta_1 - \Delta_{n+1} = \frac{2(R - kojy\bar{a} \alpha)}{R} (jy\bar{a} \alpha + jy\bar{a} 2\alpha + \dots + jy\bar{a} n\alpha)$$

$$\begin{aligned} \text{Now } \frac{2(R - kojy\bar{a} \alpha)}{R} &= \left( \frac{2jy\bar{a} \alpha}{R} \right)^2 \\ &= \left( \frac{jy\bar{a} \alpha}{R} \right)^2 \text{ approximately} \\ &= \frac{1}{225} \text{ approximately.} \end{aligned}$$

Therefore

$$jy\bar{a} (n+1)\alpha = jy\bar{a} n\alpha + jy\bar{a} \alpha - (jy\bar{a} \alpha + jy\bar{a} 2\alpha + \dots + jy\bar{a} n\alpha) / 225$$

Also

$$\begin{aligned} \Delta_{n+1} &= \Delta_n - \frac{jy\bar{a} n\alpha}{225} \\ &= \Delta_n - \frac{\Delta_1 + \Delta_2 + \dots + \Delta_n}{225}. \end{aligned}$$

Of these two interpretations the first has been given by the commentator Para-meśvara and the second by the commentators Prabhākara, Sūryadeva (b. 1191), Yallaya (1480) and Raghunātharāja (1597).

It should be observed that Āryabhaṭa I does not appear to have used this formula consistently to calculate the whole table. For as will be found from the accompanying table, certain values actually recorded by Āryabhaṭa differ from the values calcu-

TABLE

Differences $\Delta_n, n=$	Calculated according to the formula	Recorded by Āryabhaṭa	Calculated according to the modern method
1	225	225	224.856
2	224	224	223.893
3	222.005	222	221.971
4	219.018	219	219.100
5	215.045	215	215.289
6	210.089	210	210.557
7	204.156	205	204.923
8	198.245	199	198.411
9	191.36	191	191.050
10	182.512	183	182.872
11	173.694	174	173.909
12	163.245	164	164.202
13	153.196	154	153.792
14	142.512	143	142.724
15	130.876	131	131.043
16	118.294	119	118.803
17	105.745	106	106.053
18	92.289	93	92.850
19	78.88	79	79.248
20	64.527	65	65.307
21	50.240	51	51.087
22	36.014	37	36.648
23	21.849	22	22.051
24	6.752	7	7.361

lated by the formula. Probably he corrected the calculated values in those cases by comparison with the known values of the sines of  $30^\circ$ ,  $45^\circ$ ,  $60^\circ$ ; or what is much more likely employed the formula only to calculate the *Rsines* of intermediate arcs. Other plausible explanations of the discrepancy have been furnished by Krishnaswami Ayyangar<sup>1</sup> and Naraharayya.<sup>2</sup>

### *Varāhamihira and Lalla*

Varāhamihira's (d. 587) table contains the *Rsines* for every  $3^\circ 45'$  and the successive differences of the tabular *Rsines* for the radius 60.<sup>3</sup> His method of computation is this:<sup>4</sup> Starting with the known values of *Rsine*  $30^\circ$ , *Rsine*  $45^\circ$  and *Rsine*  $60^\circ$ , by the repeated and proper application of the formulae

$$\sin \frac{\theta}{2} = \frac{1}{2} \sqrt{\sin^2 \theta + \text{versin}^2 \theta}$$

$$\sin \frac{\theta}{2} = \sqrt{\frac{1}{2} \text{versin} \theta},$$

says he, the other *Rsines* may be computed. Lalla<sup>5</sup> gives a table of *Rsines* and versed *Rsines* for the radius 3438'. His method of computation is the same as that of Āryabhaṭa I and the *Sūrya-siddhānta*. He has also a shorter table of *Rsines* and their differences for intervals of  $10^\circ$  of arcs of a circle of radius 150.<sup>6</sup>

### *Brahmagupta*

Brahmagupta (628) takes the radius quite arbitrarily to be 3270. His explanation<sup>7</sup> for this departure from the usual practice is unsatisfactory.<sup>8</sup> He has, however, indicated two methods of computation.<sup>9</sup> One is *graphic* and the other *mathematical*.

*Graphic Method.* "Starting from the joint of two quadrants, mark off successively (on either directions) portions of arcs equivalent to the eighth part of a sign ( $30^\circ$ ). Join two and two of these marks by threads. Half of them (lengths of threads) will be the *Rsines*."<sup>10</sup>

*Mathematical Method.*<sup>11</sup> In this method Brahmagupta employs the trigonometrical formulae

<sup>1</sup>*JIMS*, xv, (1924), pp. 121-6.

<sup>2</sup>*Ibid*, pp. 105-13 of "Notes and Questions."

<sup>3</sup>*PSi*, iv, 6-11, 12-15.

<sup>4</sup>*Ibid*, iv, 2-5.

<sup>5</sup>*SiDVr*, ii, 1-8.

<sup>6</sup>*SiDVr*, xiii, 2-3.

<sup>7</sup>*BrSpSi*, xxi, 16.

<sup>8</sup>Datta, Bibhutibhusan, "Hindu Values of  $\pi$ ", *JASB*, N. S., Vol. 22 (1926), pp. 25-42; see particularly p. 32, footnote 1.

<sup>9</sup>His table will be found in *BrSpSi*, ii, 2-9.

<sup>10</sup>*BrSpSi*, xxi, 17.

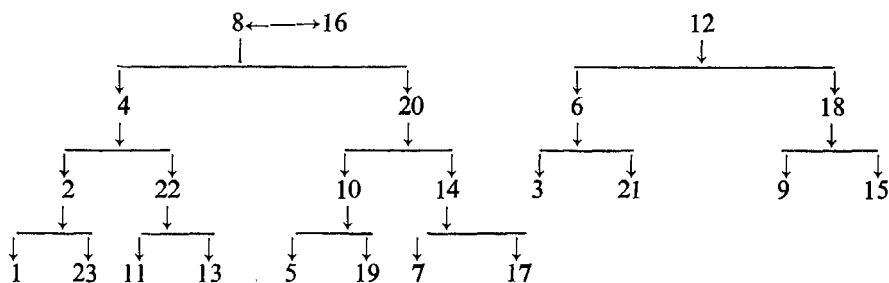
<sup>11</sup>*Ibid*, xxi, 20-21; compare also the verse 23.

$$(a) \sin \frac{\theta}{2} = \frac{1}{2} \sqrt{\sin^2 \theta + \text{versin}^2 \theta}$$

$$(b) \sin (90^\circ - \frac{\theta}{2}) = \sqrt{1 - \sin^2 \frac{\theta}{2}}$$

From the known value of the Rsine of  $8\alpha$ , that is, of  $30^\circ$ ,  $\alpha$  being equal to  $3^\circ 45'$ , we can calculate, by (a), the Rsines of  $4\alpha$ ,  $2\alpha$ ,  $\alpha$ . Then by (b) will be obtained the Rsine of  $20\alpha$ ,  $22\alpha$ ,  $23\alpha$ . Again from the first two of the latter results, we shall obtain, by (a), the Rsines of  $10\alpha$  and  $11\alpha$ ; and thence by (b) the Rsines of  $14\alpha$  and  $13\alpha$ . Continuing similar operations, we can compute the Rsines of  $5\alpha$  and  $19\alpha$ ,  $7\alpha$  and  $17\alpha$ . Again starting with the Rsine of  $12\alpha$ , we shall obtain on proceeding in the same way, successively the values of the Rsines of  $6\alpha$  and  $18\alpha$ ;  $3\alpha$  and  $21\alpha$ ;  $9\alpha$  and  $15\alpha$ . Thus the values of all the twenty-four Rsines are computed.

It is perhaps noteworthy that  $R\sin n\alpha$  is called by Brahmagupta as the  $n$ th Rsine. The successive order in which the various Rsines have been obtained above can be exhibited as follows:



Brahmagupta then observes: "In this way (can be computed) the Rsines in greater or smaller numbers, having known first the Rsines of the sixth, fourth and third parts of the circumference of the circle."<sup>1</sup> He further remarks that the Rsine of the semi-arc can be more easily calculated by the second formula of Varāhamihira.<sup>2</sup> Brahmagupta has also another table giving differences of Rsines for every  $15^\circ$  of a circle of radius 150.<sup>3</sup>

*Āryabhaṭa II and Śrīpati*

Āryabhaṭa II (950) gives the same table as that of the *Sūrya-siddhānta*.<sup>4</sup> But his method of computation is entirely different.<sup>5</sup> He takes recourse to the formulae

$$\sin \frac{1}{2} (90^\circ \pm \theta) = \sqrt{\frac{1}{2} (1 \pm \sin \theta)}$$

<sup>1</sup>*BrSpSi*, xxi. 22.

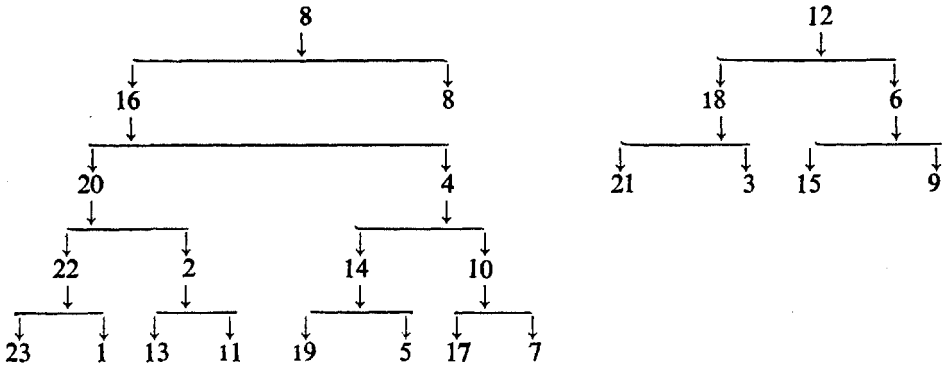
<sup>2</sup>*Ibid*, xxi. 23.

<sup>3</sup>*KK*, Part I, iii. 6; *DhGr*, 16.

<sup>4</sup>*MSi*, iii. 4-8.

<sup>5</sup>*MSi*, iii. 1-3.

Beginning with the known values of  $R\sin 30^\circ$  and  $R\sin 45^\circ$ , like Brahmagupta, the successive order in which the  $R\sin$ s will come out in the course of computation, can be best exhibited thus:



The table of Śrīpati (c. 1039) gives the  $R\sin$ s and versed  $R\sin$ s for every  $3^\circ 45'$  of a circle of radius 3415.<sup>1</sup> His *first* method of computing it is the same as the graphic method of Brahmagupta. He says:

“Place marks at the eighth parts of a sign ( $30^\circ$ ); then (starting) from the joint of two quadrants, following up these marks, join two and two of them successively by means of threads; half of them will be the  $R\sin$ s.”<sup>2</sup>

The *Second* method followed by Śrīpati is identical with the mathematical method of Brahmagupta.<sup>3</sup>

### Bhāskara II

The table of Bhāskara II (1150) contains the  $R\sin$ s and versed  $R\sin$ s as well as their differences for every  $3^\circ 45'$  of a circle of radius 3438'. He has indicated several methods of computing it. The *first* is practically the same as Brahmagupta's *graphic method*. He says:

“For computing the  $R\sin$ s, take any optional radius. On a plane ground describe a circle by means of a piece of thread equal to that radius. On it mark the cardinal points and 360 degrees; so in each quadrant of the circle there will be 90 degrees. Then divide every quadrant into as many equal parts as the number of  $R\sin$ s to be computed and put marks of these divisions. For instance, if it be required to calculate 24  $R\sin$ s, there will be 24 marks. Then beginning from any of the cardinal points, and proceeding either ways, the threads connecting the successive points will be the chords. There will be thus 24 chords. Halves of these will be the  $R\sin$ s (required). So these half-chords should be measured and the results taken as the  $R\sin$ s.”<sup>4</sup>

<sup>1</sup>*SiSe*, iii, 3-10.

<sup>2</sup>*SiSe*, xvi, 9.

<sup>3</sup>*SiSe*, xvi, 14ff.

<sup>4</sup>*SiSi*, *Gola*, v, 2-6 (Gloss).

The *second* is again a reproduction of Brahmagupta's *theoretical method*:

"When twenty-four Rsines are required (to be computed), the Rsine of 30° is the eighth element; its Rcosine is the sixteenth; and Rsin 45° is the twelfth. From these three elements, twenty-four elements can be computed in the way indicated. From the eighth we get the Rsine of its half, that is, the fourth (element), its Rcosine is the twentieth. Similarly from the fourth, the second and the twenty-second; from the second, the first and the twenty-third. In the same way from the eighth are obtained the tenth and fourteenth, fifth and nineteenth, seventh and seventeenth, eleventh and thirteenth. Again from the twelfth follow the sixth and eighteenth, third and twenty-first, ninth and fifteenth. The radius is the twenty-fourth Rsine."<sup>1</sup>

The *third* method of computing trigonometrical tables described by Bhāskara II is the same as that of Āryabhaṭa II. The speciality of this method, as also of the two following, is, says Bhāskara II, that it does not employ the versed Rsine function. As for the successive order of derivation, he points out that "from the eighth Rsine (will be obtained) the sixteenth; from the sixteenth, the fourth and the twentieth; from the fourth, the tenth and fourteenth. In this way all the rest may be deduced."<sup>2</sup>

The *fourth* method of Bhāskara II is based on the application of the formula

$$R \sin \frac{1}{2}(\theta - \phi) = \frac{1}{2} \{ (R \sin \theta - R \sin \phi)^2 + (R \cos \theta - R \cos \phi)^2 \}^{\frac{1}{2}},$$

"so that knowing any two Rsines others may be derived. For instance, let one be the fourth Rsine and the other eighth Rsine. From them is derived the second Rsine. From the second and fourth, the first; and so on."<sup>3</sup>

The *fifth* method depends on the formula

$$R \sin (45^\circ - \theta) = \sqrt{\frac{1}{2}(R \cos \theta - R \sin \theta)^2}.$$

"Thus, for instance, take the eighth Rsine; its Rcosine is the sixteenth Rsine. From these the fourth is derived; and so on."<sup>4</sup>

All the theoretical methods described above require the extraction of the square-root. So Bhāskara II propounds a new method (the *sixth*) in which that will not be necessary. It is based on the employment of the formula

$$R \cos 2\theta = R - \frac{2 (R \sin \theta)^2}{R}$$

$$\text{or } \cos 2\theta = 1 - 2 \sin^2 \theta.$$

<sup>1</sup>*SiŚi, Gola*, v. 2-6 (Gloss); xiv. 10-11 (Gloss).

<sup>2</sup>*SiŚi, Gola*, xiv. 12 (Gloss).

<sup>3</sup>*SiŚi, Gola*, xv. 13 (Gloss).

<sup>4</sup>*SiŚi, Gola*, xiv. 14 (Gloss).

But this method is defective in as much as “only certain elements of a table of Rsines can be calculated thus,”<sup>1</sup> but not the whole table. This defect is present in a sense in the previous methods, for no one of the trigonometrical formulae employed in them suffices alone for the computation of a table containing more Rsines (*vide infra*).

The *seventh* method of Bhāskara II for calculating a table of twenty-four Rsines, has been described thus:

“Multiply the Rcosine by 100 and divide by 1529; diminish the Rsine by its  $\frac{1}{467}$  part. The sum of these two results will give the next Rsine and their difference the previous Rsine. Here 225 less  $\frac{1}{7}$  is the first Rsine. And by this rule can be successively calculated the twenty-four Rsines.”<sup>2</sup>

$$jyā (n\alpha \pm \epsilon) = \left( jyā n\alpha - \frac{jyā n\alpha}{467} \right) \pm \frac{100}{1529} kojyā n\alpha,$$

where  $n=1, 2, \dots, 24$ ;  $\epsilon=3^\circ 45'$ ; and

$$jyā \epsilon = 225 - \frac{1}{7}.$$

The *rationale* of this formula is as follows:

By the Addition and Subtraction Theorems,

$$\begin{aligned} jyā (n\alpha \pm \epsilon) &= \frac{1}{R} (jyā n\alpha \cdot kojyā \epsilon \pm kojyā n\alpha \cdot jyā \epsilon), \\ &= jyā n\alpha \cdot \frac{kojyā \epsilon}{R} \pm kojyā n\alpha \cdot \frac{jyā \epsilon}{R} \end{aligned}$$

Now

$$\begin{aligned} \frac{1}{R} jyā \epsilon &= \frac{1}{3438} \left( 225 - \frac{1}{7} \right) = \frac{787}{12033} = \frac{1}{15 \cdot 289707 \dots} \\ &= \frac{100}{1528 \cdot 9707 \dots} = \frac{100}{1529} \text{ nearly} \end{aligned}$$

$$\begin{aligned} \text{and } \frac{1}{R} kojyā \epsilon &= \frac{1}{R} \sqrt{R^2 - (jyā \epsilon)^2} = \sqrt{1 - \left( \frac{jyā \epsilon}{R} \right)^2}, \\ &= \sqrt{1 - \frac{1}{233 \cdot 775 \dots}} = 1 - \frac{1}{467 \cdot 550 \dots} \end{aligned}$$

<sup>1</sup>SiŚi, Gola, xiv. 15 (Gloss).

<sup>2</sup>SiŚi, Gola, xiv. 18-20.



$$= 1 - \frac{1}{467} \text{ nearly}$$

and hence the rule. This formula is very nearly accurate. For according to the modern values

$$jyā (3^\circ 45') = 224.856 \dots$$

$$\text{Therefore } \frac{1}{R} jyā (3^\circ 45') = \frac{224.856}{3438} = \frac{1}{15.28978 \dots} = \frac{100}{15.28978 \dots}$$

Bhāskara II has indicated how to compute a table of Rsines for every  $3^\circ$  of a circle of radius 3438'. He writes:

“For instance if (it be required to compute) thirty Rsines in a quadrant, half the radius is the tenth Rsine, its Rcosine is the twentieth Rsine.  $R\sin 45^\circ$  is the fifteenth Rsine;  $R\sin 36^\circ$  is the twelfth and  $R\cos 36^\circ$  the eighteenth. The Rsine of  $18^\circ$  is the sixth and its Rcosine is the twenty-fourth. Then by the rule for deriving the Rsine of the half arc from the square-root of the sum of the squares of the Rsine and versed Rsine of an arc, as stated before, from the tenth (is derived) the fifth; its Rcosine is the twenty-fifth. In that way from the twelfth (is calculated) the sixth and twenty-fourth; from the sixth, the third and twenty-seventh; from the eighteenth, the ninth and twenty-first. These are the only elements (of the table) of Rsines which can be calculated in this way. So it has been observed that ‘only certain elements etc’. Next the formula for the Rsine of half the difference of two arcs should be employed. Let the fifth be the one Rsine and the ninth the other. From them will follow the second; its Rcosine is the twenty-eighth Rsine. From these two again by employing the (previous) rule for the Rsine of semi-arcs from the square-root of the sum of the squares of the Rsine and versed Rsine, the first and fourteenth (are obtained). The remaining fourteen Rsines can also be computed in the same way.”<sup>1</sup>

Bhāskara II has further given a rule for computing a trigonometrical table for every degree. So it is called *Pratibhāgika-jyakā-vidhi* (“The rule for the Rsine of every degree”).

“Deduct from the Rsine of any arc its 6567th part; multiply its Rcosine by 10 and then divide by 573. The sum of these two results is the next Rsine and their difference the preceding Rsine. Here the first Rsine (i.e.  $R\sin 1^\circ$ ) will be 60' and other Rsines may be successively found. Thus in a circle of radius equal to 3438', will be found 90 Rsines.”<sup>2</sup>

$$jyā (\theta \pm 1^\circ) = \left( jyā \theta - \frac{jyā \theta}{6567} \right) \pm \frac{10}{573} kojyā \theta$$

where  $\theta = 1^\circ, 2^\circ, \dots, 89^\circ$ ; given  $jyā 1^\circ = 60'$ .

<sup>1</sup>*SiŚi, Gola*, xiv. 15 (Gloss).

<sup>2</sup>*SiŚi, Gola*, xiv. 16-8.

The *rationale* of this rule can be easily found : For by the Addition and Subtraction Theorems,

$$jyā (\theta \pm 1^\circ) = \frac{1}{R} (jyā \theta. kojyā 1^\circ \pm kojyā \theta. jyā 1^\circ).$$

Now it is stated that  $R=3438'$  and  $jyā 1^\circ = 60'$ . Therefore

$$\frac{1}{R} jyā 1^\circ = \frac{60}{3438} = \frac{10}{573}$$

$$\begin{aligned} \frac{1}{R} kojyā 1^\circ &= \sqrt{1 - (jyā 1^\circ/R)^2} = \left\{ 1 - \left( \frac{10}{573} \right)^2 \right\}^{1/2} \\ &= 1 - \frac{100}{2 \times 328329} = 1 - \frac{1}{6566.58} \\ &= 1 - \frac{1}{6567} \\ &= .999847723 \dots \end{aligned}$$

The denominator wrongly appears as 6569 in Bapu Deva's edition of the *Siddhānta-Śiromaṇi*.<sup>1</sup>

The short table of Bhāskara II contains differences of *Rsines* for intervals of  $10^\circ$  in a circle of radius 120.<sup>2</sup>

### Posterior Writers

Amongst the writers posterior to Bhāskara II (1150) who have given tables of trigonometrical functions, the most notable are Mahendra Sūri (1370) and Kamalākara (1658). The latter has a table of *Rsines* and their differences for every degree of arc of a circle of radius 60, while the former gives tables of *Rsines* and versed *Rsines* together with their differences for every degree of the arc of a circle of radius 3600. Mahendra Sūri has furnished also some other tables for ready reckoning in Astronomy. It is noteworthy that Mahendra Sūri's work, *Yantrarāja*<sup>3</sup>, is admittedly based upon some Arabic work. We are informed by his commentator Malayendu Sūri, a direct disciple of the author and who wrote his commentary only 12 years after the text, that the author was the court astrologer of some potentate of the name of Firoz, who is probably the famous Sultan Firoz Shah Tughluk of Delhi (1351-88 A.D.). The illustrations chosen will agree with this date.

<sup>1</sup>The text given by Bapu Deva runs as "*Svago'ṅgeṣusaḍamśena...*". It will be "*Svāgāṅgeṣusa-ḍamśena...*".

<sup>2</sup>*SiŚi, Graha*, ii. 13.

<sup>3</sup>Mss of the text with the commentary of *Yantrarāja* are available in the libraries of India Office, London (Nos. 2906-8), Benares Sanskrit College (No. 2905), Bikaner Palace (No. 760) and also at other places. Our copy has been procured from Benares.

The following Table gives the relevant details of the various *Rsine*-Tables constructed in India from time to time.

Constructor of Table	Radius chosen	Interval taken	Sexagesimal places calculated
Author of <i>SūSi</i> <sup>1</sup>	3438'	225'	1 (minutes only)
Āryabhaṭa <sup>2</sup>	3438'	225'	1 (same Table as in <i>SūSi</i> )
Varāhamihira <sup>3</sup>	120	225'	2 (minutes and seconds)
Brahmagupta (1) <sup>4</sup>	3270	225'	1
(2) <sup>5</sup>	150	15°	1
Deva <sup>6</sup>	300	10°	1
Lalla (1) <sup>7</sup>	3438'	225'	1 (same Table as in <i>Ā</i> )
(2) <sup>8</sup>	150	10°	1
Sumati <sup>9</sup>	3438'	1°	1
Govinda Svāmi <sup>10</sup>	3437'44"19'''	225	3
Vaṭeśvara <sup>11</sup>	3437'44"	56'15"	2
Mañjula <sup>12</sup>	8°8'	30°	2 (degrees and minutes)
Āryabhaṭa II <sup>13</sup>	3438'	225'	1
Srīpati <sup>14</sup>	3415'	225'	1
Udayadivākara <sup>15</sup>	12375859''' or 3437'44"19'''	225'	1 (thirds only)
Bhāskara II <sup>16</sup> (1)	3438'	225'	1
(2)	120	10°	1
Brahmadeva <sup>17</sup>	120	15°	1
Vṛddha Vaśiṣṭha <sup>18</sup>	1000	10°	1
Malayendu Sūri <sup>19</sup>	3600	1°	2
Madanapāla <sup>20</sup>	21600	1°	2
Mādhava <sup>21</sup>	3437' 44" 48'''	225'	3
Parameśvara <sup>22</sup>	3437' 44"	225'	2
Muniśvara <sup>23</sup>	191	1°	4
Kṛṣṇa-daivajña <sup>24</sup>	500	3°	1
Kamalākara <sup>25</sup>	60	1°	5
Jagannātha Samrāṭa <sup>26</sup>	60	30'	5

<sup>1</sup>*SūSi*, ii. 17-22(a-b).

<sup>2</sup>*Ā*, i. 12.

<sup>3</sup>*PSi*, iv. 6-12.

<sup>4</sup>*BrSpSi*, ii. 2-5.

<sup>5</sup>*KK*, Part 1, iii. 6; *DhGr*, 16.

<sup>6</sup>*KR*, i. 23.

<sup>7</sup>*ŚiD Vr*, I, ii. 1-8.

<sup>8</sup>*Ibid*, xiii. 3.

<sup>9</sup>*SMT* and *SK*.

<sup>10</sup>His com. on *MBh*, iv. 22.

<sup>11</sup>*VSi*, II, i. 2-26.

<sup>12</sup>*LMā*, ii. 2(c-d).

<sup>13</sup>*MSi*, iii. 4-6(a-b).

<sup>14</sup>*SiSe*, iii. 3-6.

<sup>15</sup>His com. on *LBh*, ii. 2-3.

<sup>16</sup>*SiSi*, *Gaṇita*, ii. 3-6; 13.

<sup>17</sup>*KPr*, ii. 1.

<sup>18</sup>*VVSī*, ii. 10-11.

<sup>19</sup>*YR*, i. 5, commentary.

<sup>20</sup>His com. on *SūSi* xii. 83.

<sup>21</sup>See Nilakanṭha's com. on *Ā*, ii. 12.

<sup>22</sup>His com. on *LBh*, ii. 2(c-d)-3(a-b).

<sup>23</sup>*SiSā*, ii. 3-18.

<sup>24</sup>*KKau*, ii. 4-5.

<sup>25</sup>*SITVi*, ii. pp. 244-5 (Lucknow Edition).

<sup>26</sup>*Siddhānta-samrāt*, ii. beginning.

## 6. INTERPOLATION

*Function of any arc*

For finding the trigonometrical functions of an arc, other than those whose values have been tabulated, the Hindus generally follow the principle of proportional increase. Thus the *Sūrya-siddhānta* says:

“Divide the minutes (into which the given arc is first reduced) by 225; the quotient will indicate the number of tabular *Rsines* exceeded; (the remainder) is multiplied by the difference between the (tabular) *Rsine* exceeded and that which is still to be reached and then divide by 225. The result thus obtained should be added to the exceeded tabular *Rsine*; (the sum) will be the (required) direct *Rsine*. This rule is applicable also to the case of (determining) the versed *Rsine*.”<sup>1</sup>

Brahmagupta states:

“Divide the minutes by 225, the quotient (will indicate) the number of tabular *Rsines* (exceeded); the remainder is multiplied by the (next) difference of *Rsines* and divided by the square of 15; the result is added to the (tabular) *Rsine* corresponding to the quotient. Such (is the method) for finding the *Rsine*.”<sup>2</sup>

Such rules appear also in other astronomical works.<sup>3</sup>

The method of Mañjula (932) is very simple though it yields results only roughly approximate. He says:

“The sum of the signes (in the given arc successively) multiplied by 4, 3 and 1 will give the degrees in the *Rsines* and *Rcosines* (to be found); such are the minutes.”<sup>4</sup>

This rule though it appears to be cryptic has been fully explained by the commentators; Praśastidhara (968), Parameśvara (1430) and Yallaya (1482). We shall explain it with the help of a simple illustrative example: Suppose it is necessary to find the *Rsine* of the angle  $76^\circ 30'$ . This angle can be written as 2 signs  $16^\circ 30'$ . Now the rule says that for the first sign take 4 and for the second sign 3. For the third sign of  $30^\circ$  we are to take 1, so for a portion  $16^\circ 30'$  of that sign we should take  $(16\frac{1}{2} \times 1)/30$ . The sum of these 4, 3 and  $33/60$  will be the degrees in the required *Rsine*; and they will also be the minutes of the required *Rsine*. Therefore

<sup>1</sup>*SūSi*, ii. 31-2.

<sup>2</sup>*BrSpSi*, ii. 10; compare also *KK*, Part I, iii. 6.

<sup>3</sup>*MBh*, iv. 3-4; *LBh*, ii. 2-3; *ŚiDVr*, ii. 12; *MSi*, iii. 10 $\frac{1}{2}$ ; *SiŚi*, *Graha*, ii. 10 $\frac{1}{2}$ ; *SITV*, ii. 171; *SiSe*, iii. 15.

<sup>4</sup>“*Catustrekaghnarāyaikyaṅ doḥkoṭyoraṅśakāḥ kalāḥ*.”—*LMā*, ii. 2.

$$\begin{aligned} Jy\bar{a} (76^\circ 30') &= (4+3+33/60) \text{ degrees} + (4+3+33/60) \text{ minutes} \\ &= 7^\circ 40' 33'' \end{aligned}$$

The *rationale* of this rule which has also been given by the earlier commentators, is this: Mañjula considers the circle of reference to be of radius 488' or in sexagesimal notation 8° 8'; and his Table is very short:

Arcs	Rsines	Differences
0°	0° 0'	
		4° 4'
30°	4° 4'	
		3° 3'
60°	7° 7' <sup>1</sup>	
		1° 1'
90°	8° 8'	

To find the *Rsine* of any intermediate arc he applies the principle of proportional increase. And this at once leads to the rule.

### *Arcs of Functions*

The Hindus employed the principle of proportional increase also for the inverse problem of finding the arc which has a given trigonometrical function different from those tabulated. The *Sūrya-siddhānta* says:

“Subtract the (nearest smaller tabular) *Rsine* (from the given *Rsine*) multiply the remainder by 225 and divide by that difference (i.e. the difference corresponding to the interval in which the given *Rsine* lies); the quotient added to the number (corresponding to the tabular *Rsine* subtracted) multiplied by 225 will give the arc (required).”<sup>2</sup>

Brahmagupta writes:

“Subtract the (nearest smaller tabular) *Rsine* (from the given *Rsine*); the remainder is multiplied by 225 and divided by the (tabular) difference of *Rsines*; the

<sup>1</sup>Accurately speaking

$$jy\bar{a} 60^\circ = 488 \times \frac{\sqrt{3}}{2} = 7^\circ 2'.608\dots;$$

Mañjula takes the value to be 7° 7' obviously with the purpose of simplifying his rule.

<sup>2</sup>*SūSi*, ii, 33.

result should then be added to the product of the number corresponding to the subtracted Rsine and the square of 15: (this will be) the arc (required)."<sup>1</sup>

Similarly in other works.<sup>2</sup>

### Second Difference

The process explained above for calculating the trigonometrical functions of a given arc or the arc having given trigonometrical functions, will yield results correct only to a first degree of approximation in as much as the first difference alone of the tabular Rsines has been employed.<sup>3</sup> More accurate results will be obtained by taking into consideration also the second (and higher) differences. The earliest Hindu writer to do so was Brahmagupta. It is perhaps noteworthy that this more correct method of interpolation does not occur in his bigger work, *Brāhma-sphuṭa-siddhānta*, which was composed in 628 A.D. but in his earlier monograph *Dhyānagrahopadeśa* as well as in his later work *Khaṇḍa-Khādyaka* written in 665 A.D. These latter works, as has been stated before, contain a table of differences of Rsines for every arc of 15° in a circle of radius equal to 150. He says:

"Half the difference between the (tabular) difference passed over and that to be passed is multiplied by (residual) minutes and divided by 900; half the sum of those differences plus or minus that quotient according as it is less or greater than the (tabular) difference to be passed, will be the (corrected) value of the difference to be passed over."<sup>4</sup>

Suppose it is required to calculate the function—Rsine, Rcosine or versed Rsine—of an arc  $\alpha'$ . Let  $\alpha_1, \alpha_2, \alpha_3$  be the three consecutive values of the argument in the table such that  $\alpha_3 > \alpha_2 > \alpha_1$ .

Values of the argument $\alpha$	Values of the function $f(\alpha)$	Differences of functions
$\alpha_1$	$f_1$	
$\alpha_2$	$f_2$	$\Delta_1$
$\alpha_3$	$f_3$	$\Delta_2$

<sup>1</sup>*BrSpSi*, ii. 11; compare also *KK*, iii. 12; *DhGr*, 70.

<sup>2</sup>*MBh*, viii. 6; *SiDVr*, ii. 13; *MSi*, iii. 12; *SiSe*, iii. 16; *SiSi, Graha*, ii. 11f; *SiTVi*, ii. 172-3.

<sup>3</sup>The roughness of the result is due also to other causes. Bhāskara II observes: "As much large the radius of the circle is and into as many large number of (equal) arcs its quadrant is divided, so much accurate will be the Rsines (calculated). Otherwise (the result) will be rough (*sthūla*)".

—*SiSi, Graha*, ii. 15 (Gloss).

<sup>4</sup>"*Gatabhogyakhaṇḍakāntarodalavikalavadhācchatairnavabhīrāptyā Tadyutidalan̄ yuronam̄ bhogyādūnādhikam̄ bhogyam*" *KK*, Part 2, i.4; *DhGr*, 17. See Sengupta, P. C., "Brahmagupta on Interpolation," *BCMS*, xxii, 1931.

Now if  $\alpha_3 > \alpha' > \alpha_2$  for the calculation of  $f(\alpha')$ ,  $f_2$  will be technically called "the function exceeded,"  $f_3$  "the function to be reached,"  $\Delta_1$  "the difference passed over" and  $\Delta_2$  "the difference to be passed." Let  $\alpha' - \alpha_2 = r$ .

Now suppose that  $\alpha_3 - \alpha_2 = \alpha_2 - \alpha_1 = h$ , say. Then according to the rules stated by all Hindu astronomers,

$$f(\alpha') = f_2 + \frac{r}{h} \Delta_2;$$

which is correct up to the first order of approximation. To get more accurate results, says Brahmagupta

$$\frac{\Delta_1 + \Delta_2}{2} \pm \frac{r}{h} \left( \frac{\Delta_1 \sim \Delta_2}{2} \right)$$

should be taken as the value of "the difference to be passed", instead of  $\Delta_2$ ; the positive or negative sign being taken, according as

$$\frac{\Delta_1 + \Delta_2}{2} < \text{or} > \Delta_2.$$

Therefore, according to the method of Brahmagupta

$$f(\alpha') = f_2 + \frac{r}{h} \left\{ \frac{\Delta_1 + \Delta_2}{2} \pm \frac{r}{h} \left( \frac{\Delta_1 \sim \Delta_2}{2} \right) \right\}.$$

In the rule  $h$  is stated to be 900, as it was formulated with a view at the table of the *Khaṇḍa-Khāḍyaka*, in which the interval between the consecutive values of the argument, is  $15^\circ$  or  $900'$ . This equation can be written in the form

$$f(\alpha') = f_2 + \frac{r}{h} \Delta_2 + \frac{r}{2h} \left( 1 \pm \frac{r}{h} \right) (\Delta_1 \sim \Delta_2);$$

which agrees with the formula method of interpolation, correct up to the second degree.<sup>1</sup>

<sup>1</sup>It may be mentioned here that the formula

$$f(\alpha') = f_2 + \frac{r}{h} \Delta_1 - \frac{r}{2h} \left( 1 + \frac{r}{h} \right) (\Delta_1 - \Delta_2)$$

has been stated by Vaṭeṣara (904) in his *Siddhānta* (Ch. 2, sec. 1, vss. 65-6) and the formula

$$f(\alpha') = f_2 + \frac{r}{h} \Delta_2 + \frac{r}{2h} \left( 1 - \frac{r}{h} \right) (\Delta_1 - \Delta_2)$$

by Govindasvāmi (8th century) in his com. on *Mahā Bhāskariya* (iv. 22) and by Parameśvara (1408) in his com. on *Laghu Bhāskariya* [ii. 2(c-d)-3(a-b)]. Govindasvāmi has, however, prescribed it for the second sign only.

As has been observed by Bhāskara II<sup>1</sup> in the above formula one has to take the negative sign in calculating the *Rsine* functions and the positive sign for the versed *Rsines*. For in case of *Rsine* functions, the first difference continuously decreases as the argument increases, while contrary is the case with the versed *Rsine* functions.

Therefore the mean value of any two differences  $\left( \text{i.e. } \frac{\Delta_1 + \Delta_2}{2} \right)$  is greater than the succeeding one ( $\Delta_2$ ) in case of *Rsine* functions and less in case of versed *Rsine* functions.

Brahmagupta's method of interpolation appears also in the works of Mañjula (932) thus:

“(Find) half the sum of the tabular difference passed over and that to be passed; half their difference is multiplied by the (remaining) degrees etc. and divided by 30; half the sum minus this quotient will be the corrected value of the difference of (tabular) *Rsines* to be passed in the (calculations of the *Laghu*)—*Mānasa*.”<sup>2</sup>

The divisor is stated to be 30 in this rule, as Mañjula's table of *Rsines* contains values at intervals of 30° each.

Bhāskara II (1150) writes:

“The difference of the (tabular) difference passed over and that to be passed is multiplied by the remaining degrees and divided by 20; half the sum of the (tabular) difference passed over and that to be passed minus or plus that quotient will be the corrected value of the difference to be passed over in calculation here for *Rsines* and versed *Rsines*.”<sup>3</sup>

In formulating this rule Bhāskara II had in view a table calculated at intervals of 10°. The *rationale* of the rule has been explained by him thus:

<sup>1</sup>“...ūnam kriyate yataḥ kramajyākaraṇe khaṇḍānyapacayena vartante. Utkramajyākaraṇe tūpacayenātastatra yutamityupapannam.”—*SiŚi, Graha*, ii. 16 (Gloss).

<sup>2</sup>“Gataiśyakhāṇḍayogārddhamantarārddhena saṅguṇāt

Bhāgādeḥ khāgnilabdhanam bhogyajyā Mānase sphuṭāḥ.”

There is a bit of uncertainty about the authenticity of this verse. In the Calcutta University Collection, there are three manuscripts of the *Laghu-Mānasa* and four commentaries which contain also the text. The commentary of Praśastidhara (958) is “copied from Ms. No. B 583 and compared with other Mss. in the Oriental Library, Mysore.” That by Parameśvara (1430) is “copied from a palm leaf manuscript in Malayalam character belonging to the office of the curator for the publication of Sanskrit Mss., Trivandrum.” The source of the commentary of Yallaya (1482) is not mentioned. The above verse appears in the commentaries of Praśastidhara and Yallaya. But in the latter it has been attributed to Mallikārjuna-sūri. Now this writer flourished about 1180 A.D. Thus he is posterior to Praśastidhara by more than two centuries. So it is not possible for the latter to borrow anything from the former. I think the mistake has been made by Yallaya. The verse in question seems to me to be due in fact to Mañjula, and is more particularly from his *Bṛhat-mānasa*, which is now lost. Praśastidhara has quoted copiously from that work in his commentary of the *Laghu-mānasa* without, however, expressly mentioning it.

<sup>3</sup>*SiŚi, Graha*, ii.16.



“Half the sum of the (tabular) difference passed over and that to be passed will be the difference at the middle of those differences. But the difference to be passed is at the end of that interval to be passed. Hence proportion (should be taken) with their difference: If for an interval of  $10^\circ$ , we obtain half the difference of them, then what will be obtained for (an interval of) the remaining degrees? Thus by the rule of three, 20 will be the divisor of the product of the remaining degrees and the difference of the (tabular) difference passed over and that to be passed. By the quotient is then diminished half the sum of the (tabular) difference passed over and that to be passed; for in the calculations of Rsines the differences are in the decreasing order. But in the calculations of versed Rsines they are in the increasing order and hence the plus in this case. Thus (the rule) is proved.”

This method of interpolation has been severely criticised by Kamalākara.<sup>1</sup> But he is wrong. Muniśvara (1646) attempted to modify this method by iterating the process but his process of iteration is in correct as he has replaced the (tabular) difference to be passed by the instantaneous difference, at every stage.<sup>2</sup>

The *rationale* of the rule can be shown with the help of trigonometry to be as follows:

$$\Delta_1 = \sin \alpha_2 - \sin \alpha_1 = \sin \alpha_2 - \sin (\alpha_2 - h) = \sin \alpha_2 (1 - \cos h) + \cos \alpha_2 \sin h.$$

Since  $h$  is small we can expand  $\cos h$  and  $\sin h$  in powers of  $h$ ; then neglecting powers higher than the second, we get

$$\Delta_1 = h \cos \alpha_2 + \frac{h^2}{2} \sin \alpha_2.$$

$$\text{Similarly, } \Delta_2 = h \cos \alpha_2 - \frac{h^2}{2} \sin \alpha_2.$$

$$\text{Therefore, } \frac{\Delta_1 + \Delta_2}{2} = h \cos \alpha_2 \text{ and } \frac{\Delta_1 - \Delta_2}{2} = \frac{h^2}{2} \sin \alpha_2.$$

Now, if  $\alpha' = \alpha_2 + r$ , up to the second order of approximation, we have

$$\sin \alpha' = \sin (\alpha_2 + r) = \sin \alpha_2 \left( 1 - \frac{r^2}{2} \right) + r \cos \alpha_2,$$

Therefore,

$$\sin \alpha' = \sin \alpha_2 + \frac{r}{h} \left( \frac{\Delta_1 + \Delta_2}{2} - \frac{r}{h} \frac{\Delta_1 - \Delta_2}{2} \right).$$

<sup>1</sup> *Sitvi*, ii. 175-83.

<sup>2</sup> For Muniśvara's process of iteration, see Gupta R. C., "Muniśvara's modification of Brahmagupta's rule for second order interpolation," *IJHS*, vol. 14, no. 1, 1979, pp. 66-72.

Evidently in this case ,

$$\frac{\Delta_1 + \Delta_2}{2} > \Delta_2$$

$$\text{Hence, } \sin \alpha' = \sin \alpha_2 + \frac{r}{h} \left( \frac{\Delta_1 + \Delta_2}{2} - \frac{r}{h} \frac{\Delta_1 \sim \Delta_2}{2} \right)$$

In case of versin functions

$$\Delta_1 = \text{versin } \alpha_2 - \text{versin } (\alpha_2 - h),$$

$$= \cos (\alpha_2 - h) - \cos \alpha_2 = h \sin \alpha_2 - \frac{h^2}{2} \cos \alpha_2;$$

$$\Delta_2 = \text{versin } (\alpha_2 + h) - \text{versin } \alpha_2,$$

$$= \cos \alpha_2 - \cos (\alpha_2 + h) = h \sin \alpha_2 + \frac{h^2}{2} \cos \alpha_2.$$

$$\text{Therefore } \frac{\Delta_1 + \Delta_2}{2} < \Delta_2$$

$$\text{and } \frac{\Delta_1 + \Delta_2}{2} = h \sin \alpha_2, \quad \frac{\Delta_1 - \Delta_2}{2} = -\frac{h^2}{2} \cos \alpha_2.$$

$$\text{Now versin } \alpha' = 1 - \cos \alpha' = 1 - \cos (\alpha_2 + r),$$

$$= \text{versin } \alpha_2 + r \sin \alpha_2 + \frac{r^2}{2} \cos \alpha_2.$$

Therefore,

$$\text{versin } \alpha' = \text{versin } \alpha_2 + \frac{r}{h} \left( \frac{\Delta_1 + \Delta_2}{2} + \frac{r}{h} \cdot \frac{\Delta_1 \sim \Delta_2}{2} \right).$$

Combining these two results we get Brahmagupta's formula.

This method of interpolation has been applied also to the inverse problem of finding the arc having a given trigonometrical function.

Brahmagupta says:

“To find the arc, multiply the residue (after subtracting as many Rsines from the given quantity as possible) by 900 and divide by the difference to be passed after having determined that difference by repeated operations. The degrees in the quotient will be the arc of the residue. Subtract (as many possible) Rsines (from the given quantity), multiply the residue by 900 and divide by the (next) difference not subtracted; the quotient will be the second residue; multiply it by half the difference between the

(tabular) difference passed over and that to be passed and then divide by 900. With this quotient proceed as before for the (adjusted) value of the (tabular) difference to be passed. Repeat the same operations with the residue until the result is obtained finally.”<sup>1</sup> (By “*Rsines*” in this rule is meant “tabular *Rsine*-differences.”)

The latter portion of this rule has become rather cryptic, as all the successive operations have not been fully described. But it appears from the explanations of the commentator Bhaṭṭotpala (966) that Brahmagupta has intended the same formula as has been clearly described by Bhāskara II. The latter says:

“Subtract the (tabular) differences (as many as possible from the given value); multiply half the remainder by the difference of the (tabular) difference passed over and that to be passed, and then divide by the (tabular) difference to be passed. Half the sum of the (tabular) difference passed over and that to be passed plus or minus the quotient is the adjusted value of the (tabular) difference to be passed, whence (will follow) the arc (required).”<sup>2</sup>

He then remarks as before that the negative sign should be taken in calculating the *Rsines* and the positive sign for the versed *Rsines*.

The *rationale* of this will be clear from the previous formula on inversion. There we shall now have  $f(\alpha')$  known and  $r (= \alpha' - \alpha_2)$  as unknown.

$$f = f_2 + \frac{r}{h} \left( \frac{\Delta_1 + \Delta_2}{2} \pm \frac{r}{h} \frac{\Delta_1 - \Delta_2}{2} \right)$$

$$\text{or } \frac{r}{h} \Delta_2 = f - f_2 + \frac{r}{h} \frac{\Delta_2 - \Delta_1}{2} \mp \frac{r^2}{h^2} \frac{\Delta_1 - \Delta_2}{2}.$$

Now let us take for the first approximation, as before

$$\frac{r}{h} = \frac{f - f_2}{\Delta_2}.$$

Substituting this value of  $\frac{r}{h}$  in the neglected terms; we get for the second approximation

$$\begin{aligned} r &= (f - f_2) \frac{h}{\Delta_2} \left( 1 + \frac{\Delta_2 - \Delta_1}{2\Delta_2} \mp \frac{f - f_2}{\Delta_2^2} \frac{\Delta_1 - \Delta_2}{2} \right) \\ &= (f - f_2) \frac{h}{\Delta}, \text{ say} \end{aligned}$$

<sup>1</sup>KK, Part 2, i. 12. The printed text gives only the earlier part of the rule. The remaining portion has been taken from the text of Bhaṭṭotpala.

<sup>2</sup>SiSi Graha, ii. 17.

so that  $\Delta$  will be the adjusted value of the (tabular) difference to be passed. Then

$$\frac{1}{\Delta} = \frac{1}{\Delta_2} \left( 1 + \frac{\Delta_2 - \Delta_1}{2\Delta_2} \mp \frac{f-f_2}{\Delta_2^2} \frac{\Delta_1 - \Delta_2}{2} \right)$$

$$\text{Therefore, } \Delta = \Delta_2 \left( 1 + \frac{\Delta_2 - \Delta_1}{2\Delta_2} \mp \frac{f-f_2}{\Delta_2^2} \frac{\Delta_1 - \Delta_2}{2} \right)^{-1}$$

$$= \Delta_2 \left( 1 - \frac{\Delta_2 - \Delta_1}{2\Delta_2} \pm \frac{f-f_2}{\Delta_2^2} \frac{\Delta_1 - \Delta_2}{2} \right)$$

$$\text{or } \Delta = \frac{\Delta_1 + \Delta_2}{2} \pm \frac{f_1 - f_2}{2\Delta_2} (\Delta_1 - \Delta_2),$$

as stated in the rule. But the more accurate result by inversion would have been

$$\Delta = \frac{\Delta_1 + \Delta_2}{2} \pm \frac{f-f_2}{\Delta_1 + \Delta_2} (\Delta_1 - \Delta_2).$$

Bhāskara II was clearly aware of this. For he is found to have remarked:

“This improved formula for calculating the arc is a little rough (*sthūla*). Though rough, it has been adopted for its simplicity (*sukhārtha*). By other means such as finer calculations or repeated applications it can be made more accurate.”<sup>1</sup>

### Generalised Formula

Brahmagupta has extended his formula of interpolation so as to be applicable also to the case when the intervals between the consecutive tabular values of the argument are not equal. He says:

“Multiply the increase of the *Śighra* anomaly to be passed by the degrees of the increase of the *Śighra* equation passed over and divide by the increase of the *Śighra* anomaly passed over; the quotient is the (adjusted) increase of the *Śighra* equation in degrees. Multiply half the difference of that and the increase of the *Śighra* equation to be passed by the residue of the anomaly and divide by the increase of the *Śighra* anomaly to be passed; half the sum of these equations is decreased or increased by the (last) quotient, according as it (half the sum) is greater or less than the *Śighra* equation to be passed; the result will be corrected *Śighra* equation to be passed.”<sup>2</sup>

<sup>1</sup>“Idam dhanuḥkhaṇḍasphuṭikaraṇaṁ kiñcit sthūlam. Sthūlamapi sukhārthamangikṛtam. Anyathā bijakarmaṇā vā' sphuṭaṁ kartuṁ yujyate”—*Ibid* (Gloss).”

<sup>2</sup>Bhuktatagatiphalāṁśogunā bhogyatirbhuktatagatihṛtā labdham  
Bhuktatageḥ phalabhāgāstadbhogyaphalāntarārdhahataṁ  
Vikataṁ bhogyatihṛtaṁ labdhenonādhikāṁ phalaikyārđham  
Bhogyaphalādadhikonāṁ tadbhogyaphalaṁ sphuṭaṁ bhavati”—*KK*, Part 2, ii. 2-3.

That is, if  $\alpha_3 - \alpha_2 \neq \alpha_2 - \alpha_1$ , let  $\alpha_3 - \alpha_2 = h_2$  and  $\alpha_2 - \alpha_1 = h_1$ .

Then

$$f(\alpha') = f_2 + \frac{r}{h_2} \left\{ \left( \frac{1}{2} \frac{\Delta_1 \times h_2}{h_1} + \Delta_2 \right) \pm \frac{r}{h_2} \cdot \frac{1}{2} \left( \frac{\Delta_1 \times h_2}{h_1} - \Delta_2 \right) \right\},$$

or  $f(\alpha') = f_2 + \frac{r}{h_2} \Delta_2 + \frac{r}{2h_2} \left( 1 \pm \frac{r}{h_2} \right) \left( \frac{\Delta_1 \times h_2}{h_1} - \Delta_2 \right),$

where the upper or lower sign is to be taken according as

$$\frac{1}{2} \left( \frac{\Delta_1 \times h_2}{h_1} + \Delta_2 \right) < \text{ or } > \Delta_2.$$

### 7. SPHERICAL TRIGONOMETRY

#### *Solution of Spherical Triangles*

From the use in the treatment of astronomical problems, we find that the Hindus knew how to solve spherical triangles, of oblique as well as of right-angled varieties. They do not seem to possess a general method of solution in this matter unlike the Greeks who systematically followed the method of Ptolemy (c. 150) based on the well-known theorem of Menelaus (90). Still with the help of the properties of similar plane triangles and of the theorem of the square of the hypotenuse, they arrived at a set of accurate formulae sufficient for the purpose. As has been proved conclusively by Sengupta<sup>1</sup>, Braunmühl was wrong in supposing that in the matter of solution of spherical triangles the Hindus utilised the method of projection contained in the Analemma of Ptolemy.

#### *Right-angled Spherical Triangle*

The Hindus obtained the following formulae for the right-angled spherical triangle, right-angled at C:

- (i)  $\sin a = \sin c \sin A,$
- (ii)  $\cos c = \cos a \cos b,$
- (iii)  $\sin c \cos A = \cos a \sin b,$
- (iv)  $\sin b = \tan a \cot A,$
- (v)  $\cos A = \tan b \cot c.$

It should be particularly noted that as the tangent and cotangent functions are

<sup>1</sup>Sengupta, P. C., "Greek and Hindu Methods in Spherical Trigonometry", *Journ. Dept. Letters. Cal. Univ.*, Vol. xxi (1931).

<sup>2</sup>Braunmühl, *Geschichte*, pp. 38ff; compare also Heath, *Greek Math.*, II, p. 291.



Substituting the values in terms of trigonometrical functions, we get

$$\frac{jy\bar{a} c}{R} = \frac{nm}{kojy\bar{a} A} = \frac{jy\bar{a} a}{jy\bar{a} A} \quad (1.1)$$

$$\frac{kojy\bar{a} c}{kojy\bar{a} b} = \frac{kojy\bar{a} a}{R} = \frac{nm}{jy\bar{a} b}. \quad (2.1)$$

Hence from (1.1), we get

$$jy\bar{a} a = \frac{jy\bar{a} c \cdot jy\bar{a} A}{R},$$

which is of course equivalent to

$$\sin a = \sin c \sin A. \quad (3.i)$$

Similarly from (2.1)

$$kojy\bar{a} c = \frac{kojy\bar{a} a \cdot kojy\bar{a} b}{R}$$

$$\text{or } \cos c = \cos a \cos b. \quad (3.ii)$$

Again equating the values of  $nm$  from (1.1) and (2.1), we get

$$jy\bar{a} c \cdot kojy\bar{a} A = kojy\bar{a} a \cdot jy\bar{a} b;$$

$$\text{that is, } \sin c \cos A = \cos a \sin b, \quad (3.iii)$$

Eliminating  $\sin c$  between (3.i) and (3.iii), we get

$$\sin b = \tan a \cot A; \quad (3.iv)$$

and eliminating  $\cos a$  between (3.ii) and (3.iii), we have

$$\cos A = \tan b \cot c. \quad (3.v)$$

As an illustration of the application of the above formulae let us take the problem of determination of the Sun's right ascension ( $\alpha$ ) when the Sun's longitude ( $\lambda$ ) and declination ( $\delta$ ) are known. Let  $\gamma M$  be the equator,  $\gamma S$  the ecliptic and  $S$  the Sun. Then  $\lambda (= \gamma S)$  denotes the sun's longitude,  $\delta (= SM)$  the Sun's declination, and  $\alpha (= \gamma M)$  the Sun's right ascension. If  $\epsilon$  denotes the obliquity of the ecliptic, then by the formula (iii), we get

$$\sin \alpha = \frac{\sin \lambda \cdot \cos \epsilon}{\cos \delta}$$

This result is stated by Āryabhaṭa<sup>1</sup>, Brahmagupta<sup>2</sup> and others<sup>3</sup>. It is noteworthy that in this particular case the triangles  $Bnm$  and  $OQT$  are called technically *Krānti-Kṣetra* or "declination triangles" which shows definitely that they were actually drawn and the final result was actually obtained by the method stated above.<sup>4</sup>

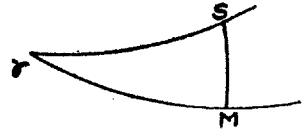


Fig. 14

### Oblique Spherical Triangle

For the solution of an oblique spherical triangle the Hindus had equivalents of the following formulae:

$$(i) \cos a = \cos b \cos c + \sin b \sin c \cos A,$$

$$(ii) \sin a \cos C = \cos c \sin b - \sin c \cos b \cos A,$$

$$(iii) \frac{\sin a}{\sin A} = \frac{\sin b}{\sin B} = \frac{\sin c}{\sin C}$$

Let  $ABC$  be a spherical triangle lying on a sphere whose centre is at  $O$ . Produce the arc  $AC$  to  $S$  and arc  $CB$  to  $Q$ , such that the arc  $CS = \text{arc } CQ = 90^\circ$ . Then  $C$  will be the pole of the great circle  $SQN$ . Join  $OA, OS$  and  $ON$ . Through  $B$  draw the small circle  $RVR'$  perpendicular to  $OA$ , intersecting the great circle  $SQN$  at  $D$  and  $D'$ . Join  $DD'$  intersecting  $SON$  at  $V$ . Let the diameter  $RVR'$  of the small circle cut  $OA$  at  $O'$ . Again through  $O$  draw the straight line  $WOE$  parallel to  $DVD'$ . Draw the great circle  $KEK'W$  parallel to the small circle  $RBR'$  and the great circle  $EAW$  perpendicular to the latter and cutting it at  $F$  and  $F'$ . From  $B$  draw  $BT$  perpendicular to  $RVR'$ ,  $BH$  to  $DVD'$  and  $BM$  to  $OQ$ . Join  $MH$  cutting  $WOE$  at  $L$ . Draw  $NY$  perpendicular to  $OA, QQ'$  and  $MM'$  to  $OS$ . Let  $BH$  cut  $FF'$  at  $G$ .

From the similar triangles  $BMH$  and  $OYN$ , we get

$$\frac{BM}{OY} = \frac{MH}{NY} = \frac{HB}{NO}; \quad (1)$$

and from the similar triangles  $OVO'$  and  $ONY$

$$\frac{OV}{ON} = \frac{VO'}{NY} = \frac{O'O}{YO}. \quad (2)$$

<sup>1</sup>*A*, iv. 25.

<sup>2</sup>*BrSpSi*, iii. 15.

<sup>3</sup>*SūSi*, iii. 41-43; *PS*, iv. 29; *SiSi, Graha*, ii. 54-5 etc.

<sup>4</sup>Compare Sen Gupta, P.C., "Papers on Hindu Mathematics and Astronomy", Part I, Calcutta, 1916, pp. 46ff; *PS*, iv. 35 (comments).

<sup>5</sup>Braunmühl, *Geschichte*, I, p. 41; Kaye, G.R., "Ancient Hindu Spherical Astronomy", *JASB*, Vol. xv (1919), pp. 153ff; P. C. Sen Gupta, *Papers etc.*, pp. 57ff.



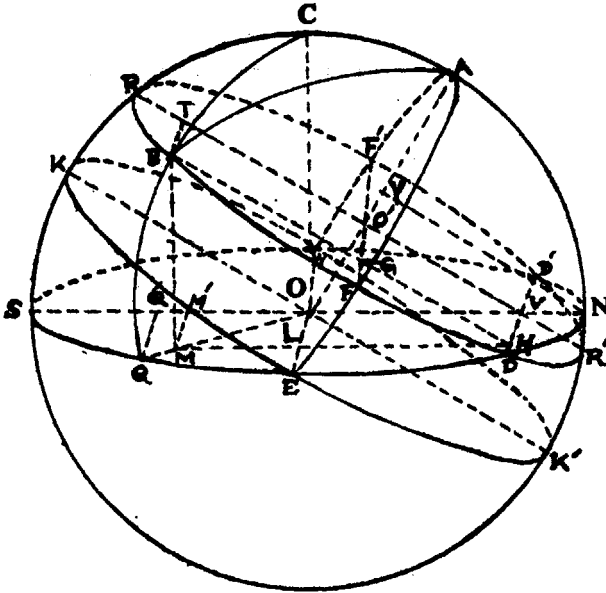


Fig. 15

Hence substituting the values

$$\frac{\text{kojyā } a}{\text{kojyā } (90^\circ - b)} = \frac{MH}{\text{jyā } (90^\circ - b)} = \frac{HB}{R}$$

$$\text{and } \frac{OV}{R} = \frac{VO'}{\text{jyā } (90^\circ - b)} = \frac{\text{kojyā } c}{\text{kojyā } (90^\circ - b)} ;$$

$$\text{whence } HB = \frac{R \text{ kojyā } a}{\text{jyā } b}, \quad MH = \frac{\text{kojyā } a \cdot \text{kojyā } b}{\text{jyā } b}, \quad (1.1)$$

$$\text{and } OV = \frac{R \text{ kojyā } c}{\text{jyā } b}, \quad O'V = \frac{\text{kojyā } b \cdot \text{kojyā } c}{\text{jyā } b} \quad (2.1)$$

Further  $O'R = \text{jyā } c,$

$$RT = \frac{\text{jyā } c \text{ utjyā } A}{R},$$

$$ML = \frac{\text{jyā } a \cdot \text{kojyā } (C - 90^\circ)}{R}$$

(3)

Now  $HB = VT = RO' + O'V - RT$ .

Therefore substituting the values of the constituent elements on either sides of equations from (1.1), (2.1) and (3), we get

$$\begin{aligned} \frac{R \text{ kojyā } a}{jyā b} &= jyā c + \frac{\text{kojyā } b. \text{ kojyā } c}{jyā b} - \frac{jyā c. \text{ utjyā } A}{R}, \\ &= \frac{\text{kojyā } b. \text{ kojyā } c}{jyā b} + \frac{jyā c \text{ kojyā } A}{R}; \end{aligned}$$

which is equivalent to

$$\cos a = \cos b \cos c + \sin b \sin c \cos A.$$

Again  $MH = ML + LH = ML + OV$ .

Therefore by (1.1), (2.1) and (3)

$$\frac{\text{kojyā } a, \text{ kojyā } b}{jyā b} = \frac{jyā a. jyā (C - 90^\circ)}{R} + \frac{R \text{ kojyā } c}{jyā b},$$

or  $R^2 \text{ kojyā } c = R \text{ kojyā } a. \text{ kojyā } b + jyā a jyā b \text{ kojyā } C$ ;

which is equivalent to

$$\cos c = \cos a \cos b + \sin a \sin b \cos C, \quad (i.1)$$

a formula similar to (i).

From the similar triangles  $OQQ'$  and  $OMM'$ , we have

$$\frac{OQ}{OM} = \frac{QQ'}{MM'} = \frac{Q'O}{M'O}. \quad (4)$$

Therefore  $OM. Q'O = OQ. OM' = OQ (MH - OV)$

$$= YN. HB - OV. R. \left[ \because \frac{MH}{YN} = \frac{HB}{NO} \right]$$

Hence  $OM. Q'O = YN (RO' + O'V - RT) - OV. R$ ;

$$\begin{aligned} \text{or } jyā a. jyā (C - 90^\circ) &= jyā (90^\circ - b) \left( jyā c + \frac{\text{kojyā } b. \text{ kojyā } c}{jyā b} - \right. \\ &\quad \left. \frac{jyā c. \text{ utjyā } A}{R} \right) - \frac{R^2 \text{ kojyā } c}{jyā b}; \end{aligned}$$

$$\text{or } -jy\bar{a} a. kojy\bar{a} C = \frac{kojy\bar{a} b. jy\bar{a} c}{R} (R - utjy\bar{a} A) - \frac{kojy\bar{a} c}{jy\bar{a} b} [R^2 - (kojy\bar{a} b)^2],$$

$$\text{or } jy\bar{a} a. kojy\bar{a} C = kojy\bar{a} c. jy\bar{a} b - jy\bar{a} c. kojy\bar{a} b. kojy\bar{a} A;$$

which is equivalent to

$$\sin a \cos C = \cos c \sin b - \sin c \cos b \cos A.$$

Since  $MM' = BT$ , we get from (4)

$$\frac{OM}{OQ} = \frac{BT}{QQ'}$$

$$\text{Hence } \frac{jy\bar{a} a}{R} = \frac{jy\bar{a} c \cdot jy\bar{a} A}{R \cdot jy\bar{a} C}$$

$$\text{or } \frac{jy\bar{a} a}{jy\bar{a} A} = \frac{jy\bar{a} c}{jy\bar{a} C} .$$

Similarly it can be proved that

$$\frac{jy\bar{a} c}{jy\bar{a} C} = \frac{jy\bar{a} b}{jy\bar{a} B} .$$

These are of course equivalent to

$$\frac{\sin a}{\sin A} = \frac{\sin b}{\sin B} = \frac{\sin c}{\sin C} .$$

As an illustration of the application of the above formulae we take up the problem of the determination of the relation between the zenith distance ( $z$ ), azimuth ( $\psi$ ) and hour angle ( $H$ ) of a heavenly body of known declination ( $\delta$ ) at a station whose terrestrial latitude is  $\phi$ . In Fig. 15 let  $NEQS$  denote the horizon,  $NACS$  the meridian circle,  $KEK'$  the equator,  $RBR'$  the diurnal circle of the heavenly body ( $B$ ),  $AFE$  the six o'clock circle,  $A$  the north pole and  $C$  the zenith. Then  $a = z$ ,  $b = 90^\circ - \phi$ ,  $c = 90^\circ - \delta$ ,  $\angle A = H$ ,  $C = \psi$ .

Substituting these values in the formulae (i), (i.1) and (ii), we obtain

$$\cos z = \sin \delta \sin \phi + \cos \delta \cos \phi \cos H$$

$$\sin \delta = \cos z \sin \phi + \sin z \cos \phi \cos \psi,$$

and

$$\sin z \cos \psi = \sin \delta \cos \phi - \cos \delta \sin \phi \cos H.$$

These equations were obtained by most of the Hindu astronomers.<sup>1</sup> It should however be made clear that the final results were arrived at by successive stages. The straight line  $DVD'$ , the line of intersection of the diurnal circle with the horizon, is technically called *udayāsta-sūtra* ("the thread through the rising and setting points"),  $BM$  is called *śaṅku* ("gnomon");  $BH$  *cheda* or *iṣṭahṛti* ("optional divisor"),  $MH$  *śaṅkutala*,  $BG$  (=the  *jyā* in the diurnal circle of the complement of the hour angle) *kalā*,  $GH$  (=the  *jyā* of the arc of the diurnal circle intercepted between the horizon and the six o'clock circle) *kujyā* or *kṣitijyā* ("earth-sine"),  $HL$ (= *jyā ED*) *agrā*,  $ML$  *bāhu*,  $OM$  *dr̥g jyā* (" *jyā* of the zenith distance"), the angle  $EAD$  *cara* ("the ascensional difference"). The existence of these technical terms proves conclusively that the Hindus actually made the constructions contemplated above. They recognise the angle  $MBH$  to be equal to the latitude of the observer's station.

The *Sūrya-siddhānta* says:

"The *Rsine* of the declination multiplied by the *palabhā* (=  $12 \tan \phi$ ) and divided by 12 gives the *kujyā* ("the earth-sine"); that multiplied by the radius and divided by the radius of the diurnal circle will give the  *jyā* whose arc will be the *cara* ("ascensional difference")."<sup>2</sup>

$$kujyā = \frac{jyā \delta \times 12 jyā \phi}{12 \times kojyā \phi}$$

$$carajyā = \frac{kujyā \times R}{kojyā \delta}$$

Again,

"The radius plus the *carajyā* in the northern hemisphere, or minus it in the southern hemisphere is called *antyā*; subtract from it the versed *Rsine* of the hour angle; (the remainder) multiplied by the radius of the diurnal circle and divided by the radius will be the *cheda*; that multiplied by the *Rsine* of the co-latitude and divided by the radius will be the *śaṅku*; subtract the square of it from the square of the radius; the square-root of the remainder will be the  *jyā* of the zenith distance."<sup>3</sup>

$$R \pm carajyā = antyā,$$

$$\frac{(antyā - utjyā H) \times kojyā \delta}{R} = cheda,$$

<sup>1</sup> *PSi*, iv. 42-4; *SūSi*, iii. 28-31, 34-5; *BrSpSi*, iii. 25-40, 54-6; *SiSi*, *Graha*, iii. 50-52; etc.

<sup>2</sup> *SūSi*, ii. 61; also *Ā*, iv. 26; *SiDVṛ*, ii. 17; *BrSpSi*, ii. 57-60; *PSi*, iv. 26-7; *SiSi*, *Graha*, ii. 48.

<sup>3</sup> *SūSi*, iii. 34-6.

$$\frac{\text{cheda} \times \text{kojyā} \phi}{R} = \text{śaṅku},$$

and  $\sqrt{R^2 - (\text{śaṅku})^2} = \text{jyā} z,$

Therefore, in the northern hemisphere,

$$\text{kojyā} z = \text{śaṅku}$$

$$= \frac{\text{kojyā} \delta \text{ kojyā} \phi}{R^2} \left( R + R \frac{\text{jyā} \delta \text{ jyā} \phi}{\text{kojyā} \delta \text{ kojyā} \phi} - \text{utjyā} H \right)$$

or  $R^2 \text{ kojyā} z = R \text{ jyā} \delta \cdot \text{jyā} \phi + \text{kojyā} \delta \cdot \text{kojyā} \phi \cdot \text{kojyā} H;$

which is of course equivalent to

$$\cos z = \sin \delta \sin \phi + \cos \delta \cos \phi \cos H.$$

Again it has been said that<sup>1</sup>

$$\text{śaṅkutala} \mp \text{bāhu} = \text{agrā},$$

the negative or positive sign being taken according as the heavenly body is in the northern or southern hemisphere. Further<sup>2</sup>

$$\text{agrā} = \frac{R \text{ jyā} \delta}{\text{kojyā} \phi} \text{ and } \text{bāhu} = - \frac{\text{jyā} z \text{ kojyā} \phi}{R};$$

Also<sup>3</sup>

$$\text{śaṅkutala} = \frac{\text{śaṅku} \times \text{jyā} \phi}{\text{kojyā} \phi}$$

Hence substituting the values

$$\begin{aligned} - \frac{\text{jyā} z \text{ kojyā} \phi}{R} &= \frac{\text{śaṅku} \times \text{jyā} \phi}{\text{kojyā} \phi} - \frac{R \text{ jyā} \delta}{\text{kojyā} \phi} \\ &= \frac{\text{jyā} \phi \text{ kojyā} \delta}{R^2} \left\{ R + R \frac{\text{jyā} \phi \cdot \text{jyā} \delta}{\text{kojyā} \phi \text{ kojyā} \delta} - \text{utjyā} H \right\} \\ &\qquad\qquad\qquad - \frac{R \text{ jyā} \delta}{\text{kojyā} \phi} \\ &= \frac{\text{jyā} \phi \text{ kojyā} \delta \text{ kojyā} H}{R^2} - \frac{\text{jyā} \delta \cdot \text{kojyā} \phi}{R}, \end{aligned}$$

<sup>1</sup>Ibid. iii. 23-4

<sup>2</sup>SūŚi, iii. 27; PSi, iv. 39; Ā, iv. 30; BrSpSi, xxi. 61.

<sup>3</sup>Ā, iv. 28, 29; BrSpSi, iii. 65, xxi. 63.

which is of course equivalent to

$$\sin z \cos \psi = \sin \delta \cos \phi - \cos \delta \sin \phi \cos H.$$

### Expansion of Trigonometrical Functions

Remarkable work on the expansion of trigonometrical functions,  $\sin \theta$ ,  $\cos \theta$ ,  $\tan^{-1}\theta$ , etc., was done in India by the astronomers of Kerala in the fourteenth, fifteenth and sixteenth centuries A.D. It will be discussed in another article which will be devoted to the "Calculus".

### ABBREVIATIONS

<i>Ā</i>	<i>Āryabhaṭīya</i>	<i>L</i>	<i>Līlāvī</i> (Ānandāśrama edition)
<i>ĀpŚiSū</i>	<i>Āpastamba-Śulba-sūtra</i>	<i>LBh</i>	<i>Laghu-Bhāskariya</i>
<i>BCMS</i>	<i>Bulletin of the Calcutta Mathematical Society</i>	<i>LMā</i>	<i>Laghu-mānasa</i>
<i>BrSpSi</i>	<i>Brāhma-sphuṭa-siddhānta</i>	<i>MBh</i>	<i>Mahā-Bhāskariya</i>
<i>DhGr</i>	<i>Dhyānagrahopadeśa</i>	<i>MSi</i>	<i>Mahā-siddhānta</i>
<i>GrL</i>	<i>Graha-lāghava</i>	<i>PSi</i>	<i>Pañca-siddhāntikā</i>
<i>GSS</i>	<i>Gaṇita-sāra-saṅgraha</i>	<i>ŚiDVṛ</i>	<i>Śiṣya-dhī-vṛddhida</i>
<i>IJHS</i>	<i>Indian Journal of History of Science</i>	<i>SiSā</i>	<i>Siddhānta-sārvabhauma</i>
<i>JIMS</i>	<i>Journal of the Indian Mathematical Society</i>	<i>SiŚe</i>	<i>Siddhānta-śekhara</i>
<i>KK</i>	<i>Khaṇḍa-khādyaka</i> (Bina Chatterjee's edition)	<i>SiŚi</i>	<i>Siddhānta-śiromaṇi</i>
<i>KKau</i>	<i>Karaṇa-kaustubha</i>	<i>SiTVi</i>	<i>Siddhānta-tattva-viveka</i>
<i>KPr</i>	<i>Karaṇa-prakāśa</i>	<i>SK</i>	<i>Sumati-karaṇa</i>
<i>KR</i>	<i>Karaṇa-ratna</i>	<i>SMT</i>	<i>Sumati-mahā-tantra</i>
		<i>SūSi</i>	<i>Sūrya-siddhānta</i>
		<i>VSi</i>	<i>Vaṭeśvara-siddhānta</i>
		<i>VVSi</i>	<i>Vṛddha-Vaṣiṣṭha-siddhānta</i>
			<i>Yantra-rāja</i>